Supplemental Material for Laurentide Ice Sheet persistence across Pleistocene interglacials

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Sediment provenance

The geochemistry and lithology of ocean sediments in both the coarse and fine fraction aids in identifying their source. In the North Atlantic, provenance tracing is complicated by the similar geologic history of the surrounding continents (Hoffman, 1988), yet some work has demonstrated that the different positioning of geologic terranes and the unique mixtures of rock types may manifest as a diagnostic signal capable of linking marine sediments to their source area(s) (Hemming et al., 2000; Hemming, 2004). Previous provenance work on the coarse fractions in or nearby cores used in this study provides strong evidence for a northeastern Canadian source for H1, H2, H3, H4, H5, and H6 on the western side of the basin as well as for H1, H2, H4, and H5 in the east (Hemming, 2004). These layers contain pronounced increases in detrital carbonate, which is perhaps the most diagnostic feature of Heinrich layers (Bond et al., 1992; Broecker et al., 1992; Hemming, 2004) and has been linked to Paleozoic basins in and near Hudson Strait using biomarker fingerprinting and sedimentation patterns (Andrews and Tedesco, 1992; Hemming, 2004; Naafs et al., 2013). H1, H2, H3, and H4 from the northwest Atlantic near Orphan Knoll (V23-14 and EW9303-GGC31), which are close to the western sites used in this study, also possess predominantly Paleoproterozoic-age grains and a smaller component of older and younger grains, which is characteristic of a Hudson Strait source (Gwiazda et al., 1996a; Hemming et al., 2000; Hemming and Hajdas, 2003; Downing and Hemming, 2012; Downing et al., 2013). Paleoproterozoic terranes can be found across Baffin Island as well as to its west and south (e.g. see Figure 1 from Downing et al., 2013). Conversely, ambient IRD between Heinrich layers from Orphan Knoll contains abundant Appalachian- and Grenville-age grains, suggesting a wider range of sources, including the southeastern LIS region (Hemming et al., 2000; Downing and Hemming, 2012). The predominantly Paleoproterozoic ages of grains across last glacial sediments and smaller components of younger and older grains suggest sediment sources came from a broad area but most likely were focused to the northwest and east of Hudson Bay. In eastern North Atlantic cores, H1, H2, H4, and H5 contain predominantly Canadian-derived IRD (Gwiazda et al., 1996a; Hemming et al., 1998) while Pb isotope data show that H3 more closely resembles ambient sediments, which are linked to a Greenland and/or a Scandinavian origin (Gwiazda et al., 1996b; Hemming et al., 1998; Snoeckx et al., 1999).

Sediment processing

We obtained a total of 456 sediment samples (10 cm³ each) from six North Atlantic sediment cores, three on the eastern side of the basin (IODP U1308 (49.8°N, 24.2°W, 3900 m), DSDP609 (49.9°N, 24.2°W, 3427 m), and BOFS-5K (50.7°N, 21.9°W, 3547)) and three on the western side (IODP U1302 (50.2°N, 45.6°W, 3560 m), IODP U1303 (50.2°N, 45.7°W, 3520 m), and EW93-02 JPC02 (48.8°N, 45.1°W, 1251 m)). The number of samples used for each Heinrich and non-Heinrich interval varied. From western sites, 22 ± 6 and 11 ± 3 samples (mean and standard deviation) were collected from each background and Heinrich layer, respectively, while 25 ± 6 and 10 ± 2 samples were collected from these intervals at eastern sites (Fig. S3-7). Samples were soaked in a light deflocculant solution of sodium metaphosphate and sieved to isolate the >63 um size fraction. Each sample was then rinsed in a 10 % HCl solution to remove carbonates. These samples were next amalgamated within Heinrich and background layers across east and west sites, producing a total of 13 samples from each side of the basin. Heinrich intervals were identified using a variety of tracers from previous work, including Ca/Sr, detrital carbonate and lithic concentrations, bulk δ^{18} O, and magnetic susceptibility (Broecker et al., 1992; Manighetti et al., 1995; Robinson et al., 1995; Stoner et al., 1996; Rosell-Mele et al., 1997; Rasmussen et al., 2003; Channell et al., 2012) (Fig. S3-7). Weakly magnetized minerals were removed via magnetic separation. Lastly, quartz was isolated through two 24-hour heated 50% HCl etches followed by repeated etches in a weak (0.125 to 0.25%) HF/HNO₃ solution. ICP-OES analysis of quartz aliquots was used to estimate sample impurities and native ²⁷Al. Samples with high impurities were re-etched until low Al concentrations (<150 ppm) were obtained.

Sample grain size

While the definition of IRD size fraction varies widely across the literature (e.g. McKay et al., 2022), we use the >63 um fraction to ensure enough material for measurement. Although the highest bottom current speeds could transport grains >63 μ m in some areas (McCave and Andrews, 2019), this size fraction is generally thought to be safe from resuspension, and we target locations where resuspension is unlikely.

U1302/03 (water depth: 3560 m and 3520 m) are located on a flank of Orphan Knoll above surrounding canyons, protecting this site from input of non-iceberg-rafted sands (e.g. turbidity currents) (Channell et al., 2012). EW93-02 JPC-02 (water depth: 1251 m) is near U1302/03 in the path of LIS icebergs and under the influence of Labrador Sea Water but above stronger Western Boundary Undercurrents, reducing the likelihood of non-IRD input in the >63 um size fraction (Hillaire-Marcel et al., 1994; Rasmussen et al., 2003). Site 609/ U1308 (water depth: 3427 m) sits within North Atlantic Deep Water but out of the path of any overflows from the Norwegian-Greenland Sea (Hodell et al., 2008), suggesting the >63 um fraction is solely from icebergs. BOFS-5K (water depth: 3547 m) is in a similar oceanographic location to Site 609/U1308 on the East Thulean Rise in the central North Atlantic.

Cosmogenic nuclide extraction

Purified Be and Al were extracted from quartz at the University of Vermont National Science Foundation Community Cosmogenic Facility using the methods outlined in Corbett et al. (2016). Samples were prepared in three batches, each containing 8-10 samples and 2-4 process blanks. We added ~0.84 g of in-house made 291 μ g/mL ⁹Be beryl carrier to each sample (totaling ~240 μ g ⁹Be), and ²⁷Al carrier (SPEX ICP standard, 1000 ppm) was added if samples contained <1500 μ g of native Al. The total ²⁷Al (quantified using ICP-OES, measured in replicate using a Yittrum internal standard) in each sample averaged 2236 ± 372 μ g. ¹⁰Be/⁹Be

ratios were measured at the Lawrence Livermore National Laboratory Center for Accelerator Mass Spectrometry and normalized to primary standard 07KNSTD3110 with an assumed ratio of 2.850 x $10^{-12.69}$ (Table S1). 26 Al/²⁷Al ratios were measured at the Purdue Rare Isotope Measurement Laboratory and normalized to KNSTD with an assumed ratio of 1.818 x $10^{-12.70}$ (Table S3).

Blanks

Twelve (¹⁰Be) and ten (²⁶Al) procedural blanks were processed alongside samples during nuclide extraction and measured to estimate background nuclides associated with laboratory processing and accelerator mass spectrometry analysis. We corrected measured sample ¹⁰Be/⁹Be and ²⁶Al/²⁷Al ratios for this contribution by subtracting the mean of all the blanks ($1.03 \pm 0.53 \times 10^{-15}$ and $7.91 \pm 0.26 \times 10^{-16}$, respectively). We quantified the blank uncertainty as the standard error of the mean and combined it with the measurement uncertainty of the sample nuclide abundances in quadrature (Tables S2 and S4).

Decay correction

Nuclide concentrations were corrected for radioactive decay on the seafloor to the time of deposition (Table S5). We used ages for Heinrich layers from Hemming (2004), and midpoints between these ages for background layers. Decay correction increases measured ²⁶Al/¹⁰Be ratios by 1-3%.

Interpretive modeling

To aid interpretation of the data, a simple 1-D model is used to simulate ice cover based on three exposure scenarios: (1) a frequent-interglacial scenario, with exposure during every Pleistocene interglacial in the LR04 δ^{18} O stack, (2) a skipped-interglacial scenario, which follows the frequent-interglacial scenario until 1 Ma but then only simulates exposure at MIS 5e, 9, and 11, and (3) a lengthened early Pleistocene interglacial scenario, similar to the skippedinterglacial scenario except that exposure periods prior to 1 Ma are twice as long as those after 1 Ma. We assess results for interglacial lengths of 1 and 10 kyr.

Results from these simulations support our conclusion that there were not extended icefree intervals across most interglacials of the past million years in sediment source regions. For instance, if there were exposure during every Pleistocene interglacial, it could only have been on the order of 1-2 kyr, otherwise ²⁶Al/¹⁰Be ratios would be higher than in most of our samples (Fig. S1 and S2, scenario 1). This result holds when limiting exposure to the few especially strong interglacials during the past million years (i.e., MIS 5e, 9, and 11) (Fig. S1 and S2, scenario 2). Simulated ²⁶Al/¹⁰Be ratios could overlap with more of our data for longer exposures (~5-10 kyr) during these few mid-to-late Pleistocene interglacials, but only if there was greater early Pleistocene exposure and much of the sediment came from very slowly eroding areas where old nuclides are preserved (Fig. S1 and S2, scenario 3). However, ¹⁰Be concentrations would be much higher than our sample values in this case.

The model calculates ¹⁰Be and ²⁶Al evolution in a bedrock column at 1 kyr time steps for various exposure and erosion scenarios over the past 8 Myr (5 Myr of exposure to initialize conditions + 3 Myr of Pleistocene ice cover change). For each model run, nuclide production is simulated during intervals of exposure, while nuclide production ceases and erosion occurs during intervals of ice cover. We assume sea-level, high-latitude production rates, including production by muons, calculated using the MATLAB implementation in Balco et al., (2008) of

the method of Heisinger et al. (2002) and a 26 Al/ 10 Be surface production ratio of 7.3 (Corbett et al., 2017). Nuclides decay continuously throughout model runs. To incorporate erosion, the "top" of the bedrock column is redefined as the eroded depth at each time step. For instance, an erosion rate of 10 mm/kyr would remove the upper 10 mm of bedrock over one timestep, and the following timestep would start with the concentration at 10 mm depth as the new surface. The model uses the following equation to calculate concentrations in the bedrock column:

$$N(t,z) = N(t-1,z) \cdot e^{\frac{((-\Delta t \ln 2)/t_1)}{2}} + P(z) \cdot \Delta t (1 - C_{ice})$$

Where N(t, z) is the nuclide (¹⁰Be or ²⁶Al) concentration, which changes with both depth (z) and time (t), N(t-1, z) is the nuclide concentration as a function of the previous time step (t-1) and depth, Δt is the time step length (1 kyr), t_{1/2} is the nuclide half-life, P(z) is the nuclide production rate as a function of depth, and C_{ice} is 1 during ice cover and 0 during exposure. Prior to 3 Ma, constant production of both nuclides and erosion at 20 mm/kyr occurs. We run each exposure scenario over a range of erosion rates (1, 10, and 100 mm/kyr) to assess the sensitivity of simulated ¹⁰Be and ²⁶Al concentrations to erosion.

To explore the potential impact of nonstationary erosion rates, we rerun exposure scenario 1 from above through three additional simulations with accelerating erosion over the Pleistocene of different baselines and magnitudes of acceleration (Fig. S8). These simulations suggest 10-kyr exposure events every interglacial could yield ²⁶Al/¹⁰Be ratios as low as we measured, but only if erosion rates accelerated one to two orders of magnitude over the Pleistocene. We are unaware of long-term erosion rate reconstructions for the LIS, but a recent reconstruction for Eurasian ice sheets suggests erosion rate changes were much smaller and in fact decreased from the mid to late Pleistocene (Lien et al., 2022). Furthermore, simulated ¹⁰Be concentrations in these accelerating erosion simulations are an order of magnitude lower than the concentrations we measured in IRD. We therefore consider accelerating erosion an unlikely explanation for the low ²⁶Al/¹⁰Be ratios we measured, and favor quasi-persistent ice cover over the past million years as the most parsimonious interpretation of the dataset.

To assess whether nuclide production through thin ice cover during glacials might yield low ²⁶Al/¹⁰Be ratios despite frequent 10-kyr interglacial exposures, we rerun scenario 1 from above but simulate nuclide production by muons through 50 m of ice when covered (Fig. S9). Simulated ²⁶Al/¹⁰Be ratios are higher than in our default simulation with no production through ice, discounting this explanation. The MATLAB code used for forward model simulations is available at https://github.com/danielleeleblanc/LeBlanc-et-al-2022-Unpublished.



Fig. S1. The ternary diagrams show how (A) ¹⁰Be concentrations and (B) ²⁶Al/¹⁰Be ratios mix between three end members as a function of the areal fraction of sediment source areas each occupies. End members are simulated values from Fig. 3 using benthic δ^{18} O threshold of 3.7 ‰ for 1, 10, and 100 mm/kyr glacial erosion rates. See Bierman et al., 2016 for mixing equations.



Fig. S2. Modeled cosmogenic nuclides for 1-kyr interglacial exposures. Simulated ¹⁰Be concentrations (left) and ²⁶Al/¹⁰Be ratios (right) over the past 3 Myr for three ice cover scenarios (top) for 1-kyr interglacial exposures using three different erosion rates: 1, 10, and 100 mm/kyr. Scenario 1 has exposure during every Pleistocene interglacial. Scenario 2 features exposure during every interglacial until 1 Ma, but only during select interglacials thereafter (MIS 1, 5e, 9, and 11). Scenario 3 is the same as scenario 2 but interglacial exposures are twice as long before 1 Ma as after (i.e., scenarios with 1-kyr exposure after 1 Ma have 2-kyr exposure before 1 Ma). The LR04 benthic oxygen isotope stack is shown (top) as a proxy for global ice volume³⁸. The mid-Pleistocene transition (MPT) marks the shift from 41-kyr to ~100-kyr glacial cycles. Boxes from 64-14 ka in lower panels show the range of our measured values.



Fig. S3. Modeled cosmogenic nuclides for 10-kyr interglacial exposures. The same as Fig. S1 but with 10-kyr interglacial exposures.



Fig. S4. Identification of Heinrich layers in U1302/3. Peaks in Ca/Sr reflect detrital carbonaterich Heinrich layers (Channell et al., 2012; Channell and Hodell, 2013). Black dots at the top indicate the locations of our samples.



Fig. S5. Identification of Heinrich layers in BOFS-5K. Peaks in magnetic susceptibility (yellow) and lithics concentration (green) reflect IRD-rich Heinrich layers (Manighetti et al., 1995; Robinson et al., 1995; Rosell-Mele et al., 1997). Black dots at the top indicate the locations of our samples.



Fig. S6. Identification of Heinrich layers in EW9302-2JPC. Peaks in detrital carbonate (blue) and IRD concentration (purple) reflect Heinrich layers (Rasmussen et al., 2003; Marcott et al., 2011). Black dots at the top indicate the locations of our samples.



Fig. S7. Identification of Heinrich layers in DSDP-609. Peaks in lithics concentration (green) and *N. pachyderma* (s.) δ^{18} O (peach) depletions reflect IRD and freshwater pulses associated with Heinrich events (Broecker et al., 1992). Black dots at the top indicate the locations of our samples.



Fig. S8. Identification of Heinrich layers in U1308. Ca/Sr (red) peaks and bulk carbonate δ^{18} O (blue) depletions reflect detrital carbonate-rich Heinrich layers (Hodell et al., 2017). Black dots at the top indicate the locations of our samples.



Fig. S9. Modeled cosmogenic nuclides for 1-and 10-kyr interglacial exposures with accelerating erosion. Simulated ¹⁰Be concentrations (left) and ²⁶Al/¹⁰Be ratios (right) over the past 3 Myr for ice cover scenario 1 shown at the top (same as in Fig. S1) with 1- and 10-kyr interglacial exposure using three different accelerating erosion scenarios (A, B, and C on top plot). Boxes from 65-14 ka show the range of our measured values. While erosion scenarios B and C yield ²⁶Al/¹⁰Be ratios as low as measured, we consider these scenarios unlikely because they imply subglacial erosion was negligible in the early Pleistocene and then increased orders of magnitude in the late Pleistocene.



Fig. S10. The same as in Fig. S2 for ice cover scenario 1 (colored lines), but also showing results including production by muons through 50 m of ice during glacials (black).

| Sample Name | Quartz Mass (g) | Mass of ⁹ Be Added (µg)* | AMS Cathode Number | Uncorrected ¹⁰ Be/ ⁹ Be Ratio** | Uncorrected ¹⁰ Be/ ⁹ Be Ratio Uncertainty** | Background- Corrected ¹⁰ Be/ ⁹ Be Ratio | Background- Corrected ¹⁰ Be/ ⁹ Be Ratio 1σ Uncertainty | ¹⁰ Be Concentration (atoms g ⁻¹) | ¹⁰ Be Concentration 1σ Uncertainty (atoms g ⁻¹) | |
|----------------|-----------------------|--|--------------------------|---|---|--|--|---|--|------|
| EB0 | 4.1419 | 243.2 | BE48152 | 6.784E-15 | 4.375E-16 | 5.903E-15 | 5.746E-16 | 2.316E+04 | 2.255E+03 | |
| EB2 | 4.9050 | 243.1 | BE48153 | 9.555E-15 | 4.931E-16 | 8.674E-15 | 6.180E-16 | 2.872E+04 | 2.046E+03 | |
| EH3 | 4.2999 | 242.9 | BE48154 | 1.017E-14 | 6.273E-16 | 9.287E-15 | 7.296E-16 | 3.505E+04 | 2.754E+03 | |
| EH6 | 4.7744 | 242.1 | BE48155 | 1.068E-14 | 6.484E-16 | 9.797E-15 | 7.478E-16 | 3.320E+04 | 2.534E+03 | ch 3 |
| EB6 | 4.9062 | 242.2 | BE48158 | 1.196E-14 | 4.996E-16 | 1.108E-14 | 6.232E-16 | 3.655E+04 | 2.056E+03 | Bat |
| WB0 | 5.8679 | 242.7 | BE48159 | 3.607E-15 | 3.528E-16 | 2.726E-15 | 5.131E-16 | 7.536E+03 | 1.418E+03 | |
| WB3 | 5.9715 | 242.8 | BE48160 | 5.959E-15 | 4.463E-16 | 5.079E-15 | 5.814E-16 | 1.380E+04 | 1.580E+03 | |
| WB4 | 4.4646 | 241.9 | BE48161 | 5.290E-15 | 5.067E-16 | 4.409E-15 | 6.290E-16 | 1.597E+04 | 2.278E+03 | |
| EB3 | 13.5617 | 240.8 | BE47420 | 3.779E-14 | 9.659E-16 | 3.632E-14 | 1.172E-15 | 4.310E+04 | 1.390E+03 | |
| EB4 | 13.6682 | 241.8 | BE47422 | 3.777E-14 | 1.134E-15 | 3.630E-14 | 1.314E-15 | 4.292E+04 | 1.553E+03 | |
| EB5 | 4.9141 | 242.0 | BE47423 | 9.790E-15 | 9.795E-16 | 8.323E-15 | 1.183E-15 | 2.739E+04 | 3.893E+03 | |
| WB2 | 4.1814 | 240.3 | BE47425 | 4.140E-15 | 3.121E-16 | 2.673E-15 | 7.331E-16 | 1.027E+04 | 2.816E+03 | ch 2 |
| WB5 | 4.3290 | 241.6 | BE47426 | 6.240E-15 | 4.722E-16 | 4.774E-15 | 8.143E-16 | 1.780E+04 | 3.037E+03 | Bat |
| WB6 | 9.4252 | 241.9 | BE47428 | 1.019E-14 | 7.778E-16 | 8.727E-15 | 1.022E-15 | 1.497E+04 | 1.753E+03 | |
| WH3 | 6.7512 | 241.5 | BE47429 | 5.513E-15 | 4.942E-16 | 4.047E-15 | 8.272E-16 | 9.674E+03 | 1.978E+03 | |
| WH6 | 8.1019 | 240.4 | BE47431 | 8.867E-15 | 4.990E-16 | 7.400E-15 | 8.301E-16 | 1.468E+04 | 1.646E+03 | |
| EH1 | 8.2163 | 242.7 | BE47239 | 5.831E-15 | 5.612E-16 | 5.092E-15 | 6.119E-16 | 1.005E+04 | 1.208E+03 | |
| WH1 | 8.7254 | 242.2 | BE47240 | 4.622E-15 | 5.207E-16 | 3.883E-15 | 5.750E-16 | 7.201E+03 | 1.066E+03 | |
| EB1 | 6.8760 | 242.5 | BE47242 | 9.713E-15 | 5.985E-16 | 8.974E-15 | 6.463E-16 | 2.115E+04 | 1.523E+03 | |
| WB1 | 9.4924 | 241.7 | BE47243 | 5.466E-15 | 4.418E-16 | 4.727E-15 | 5.046E-16 | 8.042E+03 | 8.585E+02 | |
| EH2 | 9.0450 | 241.4 | BE47244 | 6.727E-15 | 4.343E-16 | 5.988E-15 | 4.981E-16 | 1.068E+04 | 8.882E+02 | ch 1 |
| WH2 | 6.0834 | 242.4 | BE47245 | 3.815E-15 | 3.468E-16 | 3.076E-15 | 4.240E-16 | 8.191E+03 | 1.129E+03 | Bat |
| EH4 | 12.0356 | 242.2 | BE47246 | 1.177E-14 | 5.489E-16 | 1.103E-14 | 6.006E-16 | 1.484E+04 | 8.078E+02 | |
| WH4 | 12.9206 | 242.7 | BE47247 | 9.624E-15 | 5.729E-16 | 8.885E-15 | 6.226E-16 | 1.115E+04 | 7.814E+02 | |
| EH5 | 7.2881 | 241.0 | BE47249 | 7.123E-15 | 3.967E-16 | 6.384E-15 | 4.657E-16 | 1.411E+04 | 1.029E+03 | |
| WH5 | 14.5706 | 241.3 | BE47250 | 1.229E-14 | 7.775E-16 | 1.155E-14 | 8.148E-16 | 1.278E+04 | 9.017E+02 | |

 $^{*9}\mbox{Be}$ was added through a beryl carrier made at University of Vermont with a concentration of 291 μg mL $^{-1}$.

**Isotopic analysis was conducted at Lawrence Livermore National Laboratory; ratios were normalized against standard 07KNSTD3110 with an assumed ratio of 2850 x 10⁻¹⁵ ⁶⁹.

Table S1. ¹⁰Be measurements in North Atlantic ice-rafted debris. All western and eastern Heinrich layers are indicated by 'WH' and 'EH', respectively. Each number refers to the corresponding Heinrich interval. Intervening background intervals are labelled as 'WB' and 'EB' for west and east sides of the North Atlantic, respectively. Numbers for intervening layers start at 0, indicating samples before H1, and increase going down hole. Samples were processed in three batches (shown on right).

| Blank Name | UVM Batch Number | Cathode Number | Be Analysis Date | From AMS: ¹⁰ Be/ ⁹ Be Ratio | From AMS: ¹⁰ Be/ ⁹ Be Ratio 1σ Uncertainty | |
|------------|---------------------|-------------------|---------------------|--|--|------|
| BLK | 674 | BE48151 | 2/26/20 | 1.416E-15 | 2.363E-16 | |
| BLKX | 674 | BE48156 | 2/26/20 | 8.465E-16 | 1.804E-16 | ch 3 |
| BLKXX | 674 | BE48157 | 2/26/20 | 6.623E-16 | 1.562E-16 | Bat |
| BLKXXX | 674 | BE48162 | 2/26/20 | 5.968E-16 | 1.656E-16 | |
| BLK | 669 | BE47421 | 11/20/19 | 1.293E-15 | 3.138E-16 | |
| BLKX | 669 | BE47424 | 11/20/19 | 2.364E-15 | 3.451E-16 | ch 2 |
| BLKXX | 669 | BE47427 | 11/20/19 | 7.706E-16 | 1.989E-16 | Bate |
| BLKXXX | 669 | BE47430 | 11/20/19 | 1.440E-15 | 3.069E-16 | |
| BLK | 662 | BE47241 | 9/19/19 | 9.137E-16 | 1.867E-16 | |
| BLKX | 662 | BE47248 | 9/19/19 | 4.429E-16 | 1.337E-16 | sh 1 |
| BLK | 664 | BE47254 | 9/19/19 | 9.622E-16 | 1.853E-16 | Bate |
| BLKX | 664 | BE47259 | 9/19/19 | 6.370E-16 | 1.425E-16 | |
| | | | AVERAGE | 1.029E-15 | | |
| | | | STDEV | 5.3142E-16 |] | |

 Table S2. ¹⁰Be blank values. Blanks were processed in three batches (shown on right).

| Sample Name | Quartz Mass (g) | Total ²⁷ Al Quantified by ICP- OES (µg)* | AMS Cathode Number | Uncorrected ²⁶ Al/ ²⁷ Al Ratio** | Uncorrected ²⁶ Al/ ²⁷ Al Ratio Uncertainty** | Background- Corrected ²⁶ Al/ ²⁷ Al Ratio | Background- Corrected ²⁶ Al/ ²⁷ Al Ratio Uncertainty | ²⁶ Al Concentration (atoms g ⁻¹) | ²⁶ Al Concentration Uncertainty (atoms g ⁻¹) | |
|--|-----------------------|--|--------------------------|--|--|---|--|---|--|------|
| EB0 | 4.1419 | 2029 | 160449 | 1.168E-14 | 1.134E-15 | 1.089E-14 | 1.164E-15 | 1.190E+05 | 1.273E+04 | |
| EB2 | 4.9050 | 2193 | 160450 | 1.690E-14 | 1.429E-15 | 1.611E-14 | 1.453E-15 | 1.607E+05 | 1.450E+04 | |
| EH3 | 4.2999 | 2009 | 160451 | 1.965E-14 | 1.667E-15 | 1.886E-14 | 1.688E-15 | 1.967E+05 | 1.760E+04 | |
| EH6 | 4.7744 | 2088 | 160452 | 1.970E-14 | 1.611E-15 | 1.891E-14 | 1.632E-15 | 1.846E+05 | 1.593E+04 | ch 3 |
| EB6 | 4.9062 | 2348 | 160455 | 2.004E-14 | 1.643E-15 | 1.924E-14 | 1.664E-15 | 2.056E+05 | 1.778E+04 | Batc |
| WB0 | 5.8679 | 3416 | 160456 | 3.442E-15 | 7.257E-16 | 2.651E-15 | 7.717E-16 | 3.444E+04 | 1.003E+04 | 1 |
| WB3 | 5.9715 | 2528 | 160457 | 6.132E-15 | 8.844E-16 | 5.341E-15 | 9.226E-16 | 5.046E+04 | 8.716E+03 | |
| WB4 | 4.4646 | 2066 | 160458 | 8.232E-15 | 9.819E-16 | 7.441E-15 | 1.016E-15 | 7.685E+04 | 1.050E+04 | |
| EB3 | 13.5617 | 2137 | 160436 | 6.368E-14 | 4.317E-15 | 6.289E-14 | 4.325E-15 | 2.212E+05 | 1.521E+04 | |
| EB4 | 13.6682 | 2079 | 160438 | 5.436E-14 | 2.658E-15 | 5.357E-14 | 2.671E-15 | 1.818E+05 | 9.067E+03 | |
| EB5 | 4.9141 | 1953 | 160439 | 1.713E-14 | 1.510E-15 | 1.634E-14 | 1.532E-15 | 1.449E+05 | 1.359E+04 | 1 |
| WB2 | 4.1814 | 1930 | 160441 | 5.693E-15 | 8.352E-16 | 4.902E-15 | 8.755E-16 | 5.049E+04 | 9.018E+03 | sh 2 |
| WB5 | 4.3290 | 2151 | 160442 | 9.071E-15 | 1.045E-15 | 8.280E-15 | 1.077E-15 | 9.183E+04 | 1.194E+04 | Bate |
| WB6 | 9.4252 | 2377 | 160444 | 1.080E-14 | 1.089E-15 | 1.001E-14 | 1.120E-15 | 5.633E+04 | 6.306E+03 | 1 |
| WH3 | 6.7512 | 2412 | 160445 | 7.287E-15 | 1.058E-15 | 6.495E-15 | 1.091E-15 | 5.179E+04 | 8.696E+03 | |
| WH6 | 8.1019 | 2261 | 160447 | 9.476E-15 | 1.142E-15 | 8.684E-15 | 1.172E-15 | 5.410E+04 | 7.299E+03 | 1 |
| EH1 | 8.2163 | 1831 | 160424 | 1.111E-14 | 1.310E-15 | 1.032E-14 | 1.336E-15 | 5.134E+04 | 6.649E+03 | |
| WH1 | 8.7254 | 2119 | 160425 | 4.887E-15 | 8.539E-16 | 4.096E-15 | 8.933E-16 | 2.220E+04 | 4.841E+03 | |
| EB1 | 6.8760 | 1854 | 160427 | 1.792E-14 | 1.969E-15 | 1.713E-14 | 1.987E-15 | 1.031E+05 | 1.196E+04 | |
| WB1 | 9.4924 | 2466 | 160428 | 7.332E-15 | 9.514E-16 | 6.540E-15 | 9.870E-16 | 3.792E+04 | 5.722E+03 | |
| EH2 | 9.0450 | 2387 | 160429 | 7.419E-15 | 1.377E-15 | 6.628E-15 | 1.402E-15 | 3.903E+04 | 8.257E+03 | sh 1 |
| WH2 | 6.0834 | 2081 | 160430 | 5.859E-15 | 8.174E-16 | 5.068E-15 | 8.585E-16 | 3.868E+04 | 6.553E+03 | Batc |
| EH4 | 12.0356 | 2827 | 160431 | 1.418E-14 | 1.296E-15 | 1.338E-14 | 1.322E-15 | 7.016E+04 | 6.930E+03 | |
| WH4 | 12.9206 | 2897 | 160432 | 1.312E-14 | 1.374E-15 | 1.233E-14 | 1.399E-15 | 6.168E+04 | 6.998E+03 | |
| EH5 | 7.2881 | 1682 | 160434 | 1.590E-14 | 1.719E-15 | 1.510E-14 | 1.739E-15 | 7.781E+04 | 8.958E+03 | |
| WH5 | 14.5706 | 2016 | 160435 | 2.138E-14 | 1.563E-15 | 2.059E-14 | 1.585E-15 | 6.357E+04 | 4.894E+03 | |
| * ²⁷ Al was added only to samples with insufficient total Al through commercial SPEX ICP standard with a concentration of 1000 μg mL ⁻¹ . The total here reflects the sum of Al added through carrier and native Al in quartz. | | | | | | | | | | |
| **Isotopic analysis was conducted at PRIME Laboratory; ratios were normalized against standard KNSTD with an assumed ratio of 1.818 x 10 ^{-12 70} . | | | | | | | | | | |

Table S3. ²⁶Al measurements in North Atlantic ice-rafted debris. All western and eastern Heinrich layers are indicated by 'WH' and 'EH', respectively. Each number refers to the corresponding Heinrich interval. Intervening background intervals are labelled as 'WB' and 'EB' for west and east sides of the North Atlantic, respectively. Numbers for intervening layers start at 0, indicating samples before H1, and increase going down hole. Samples were processed in three batches (shown on right).

| Blank Name | UVM Batch Number | Cathode Number | Al Analysis Date | From AMS: ²⁶ Al/ ²⁷ Al Ratio | From AMS: ²⁶ Al/ ²⁷ Al Ratio Uncertainty | |
|------------|---------------------|-------------------|---------------------|---|--|------|
| BLK | 674 | 160448 | 5/14/21 | 9.834E-16 | 3.720E-16 | |
| BLKX | 674 | 160453 | 5/14/21 | 1.009E-15 | 3.824E-16 | ch 3 |
| BLKXX | 674 | 160454 | 5/14/21 | 5.103E-16 | 2.553E-16 | Bat |
| BLKXXX | 674 | 160459 | 5/14/21 | 1.220E-15 | 4.074E-16 | |
| BLK | 669 | 160437 | 5/14/21 | 3.327E-16 | 3.186E-16 | |
| BLKX | 669 | 160440 | 5/14/21 | 7.316E-16 | 3.275E-16 | ch 2 |
| BLKXX | 669 | 160443 | 5/14/21 | 5.942E-16 | 2.659E-16 | Bate |
| BLKXXX | 669 | 160446 | 5/14/21 | 9.320E-16 | 3.525E-16 | |
| BLK | 662 | 160426 | 5/14/21 | 7.947E-16 | 4.279E-16 | ch 1 |
| BLKX | 662 | 160433 | 5/14/21 | 8.046E-16 | 3.782E-16 | Bate |
| | | | AVERAGE | 7.91E-16 | | |
| | | | STDEV | 2.62E-16 | | |

 Table S4. ²⁶Al blank values. Blanks were processed in three batches (shown on the right).

| Sample Name | Age (ka) | Decay-Corrected ¹⁰ Be Concentration (atoms g ⁻¹) | Decay-Corrected ¹⁰ Be Concentration 1σ Uncertainty (atoms g ⁻¹) | Decay- Corrected ²⁶ Al Concentration (atoms g ⁻¹) | Decay-Corrected ²⁶ Al Concentration 1σ Uncertainty (atoms g ⁻¹) | Decay- Correct ed ²⁶ Al/ ¹⁰ Be | Decay- Corrected ²⁶ Al/ ¹⁰ Be 1σ Uncertainty |
|-------------|-------------|--|--|---|--|---|---|
| EB0 | 14 | 2.28E+04 | 2.72E+03 | 1.21E+05 | 1.29E+04 | 5.30 | 0.85 |
| EH1 | 16.8 | 9.56E+03 | 1.54E+03 | 5.21E+04 | 6.76E+03 | 5.45 | 1.13 |
| EB1 | 20.4 | 2.07E+04 | 1.91E+03 | 1.05E+05 | 1.22E+04 | 5.07 | 0.75 |
| EH2 | 24 | 1.03E+04 | 1.23E+03 | 3.99E+04 | 8.46E+03 | 3.87 | 0.94 |
| EB2 | 27.5 | 2.86E+04 | 2.43E+03 | 1.65E+05 | 1.49E+04 | 5.78 | 0.72 |
| EH3 | 31 | 3.50E+04 | 3.15E+03 | 2.03E+05 | 1.81E+04 | 5.80 | 0.73 |
| EB3 | 34.5 | 4.44E+04 | 1.33E+03 | 2.29E+05 | 1.57E+04 | 5.15 | 0.39 |
| EH4 | 38 | 1.47E+04 | 1.05E+03 | 7.29E+04 | 7.19E+03 | 4.96 | 0.61 |
| EB4 | 41.5 | 4.43E+04 | 1.51E+03 | 1.90E+05 | 9.45E+03 | 4.28 | 0.26 |
| EH5 | 45 | 1.38E+04 | 1.50E+03 | 8.13E+04 | 9.36E+03 | 5.89 | 0.93 |
| EB5 | 52.5 | 2.96E+04 | 3.77E+03 | 1.53E+05 | 1.43E+04 | 5.16 | 0.82 |
| EH6 | 60 | 3.37E+04 | 2.93E+03 | 1.96E+05 | 1.69E+04 | 5.82 | 0.71 |
| EB6 | 65 | 3.73E+04 | 2.49E+03 | 2.19E+05 | 1.90E+04 | 5.89 | 0.64 |
| WB0 | 14 | 7.18E+03 | 1.77E+03 | 3.49E+04 | 1.01E+04 | 4.86 | 1.85 |
| WH1 | 16.8 | 6.72E+03 | 1.39E+03 | 2.26E+04 | 4.92E+03 | 3.36 | 1.01 |
| WB1 | 20.4 | 7.63E+03 | 1.19E+03 | 3.87E+04 | 5.84E+03 | 5.07 | 1.10 |
| WH2 | 24 | 7.51E+03 | 1.71E+03 | 3.96E+04 | 6.71E+03 | 5.28 | 1.50 |
| WB2 | 27.5 | 1.21E+04 | 2.40E+03 | 5.19E+04 | 9.27E+03 | 4.30 | 1.15 |
| WH3 | 31 | 1.09E+04 | 1.76E+03 | 5.34E+04 | 8.97E+03 | 4.91 | 1.15 |
| WB3 | 34.5 | 1.36E+04 | 1.92E+03 | 5.22E+04 | 9.02E+03 | 3.83 | 0.85 |
| WH4 | 38 | 1.10E+04 | 1.00E+03 | 6.40E+04 | 7.26E+03 | 5.82 | 0.85 |
| WB4 | 41.5 | 1.57E+04 | 2.72E+03 | 8.00E+04 | 1.09E+04 | 5.09 | 1.12 |
| WH5 | 45 | 1.28E+04 | 1.06E+03 | 6.65E+04 | 5.11E+03 | 5.20 | 0.59 |
| WB5 | 52.5 | 1.99E+04 | 2.72E+03 | 9.66E+04 | 1.25E+04 | 4.85 | 0.91 |
| WH6 | 60 | 1.60E+04 | 1.49E+03 | 5.74E+04 | 7.74E+03 | 3.59 | 0.59 |
| WB6 | 65 | 1.62E+04 | 1.67E+03 | 6.00E+04 | 6.72E+03 | 3.70 | 0.56 |

Table S5. Decay-corrected ²⁶Al and ¹⁰Be concentrations and ratios in North Atlantic ice-rafted debris. Measured ¹⁰Be concentrations (from Table S1) and ²⁶Al concentrations (from Table S3) are shown here corrected for radioactive decay on the seafloor to the time of deposition.

Supplemental Materials References

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