

# GSA DATA REPOSITORY 2014131

## Late Holocene fluctuations of Qori Kalis outlet glacier, Quelccaya Ice Cap, Peruvian Andes

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### DATA REPOSITORY FIGURES AND TABLES

Figures DR1-DR10 provide descriptions of all  $^{10}\text{Be}$  ages and interpreted moraine ages as well as photos of boulders sampled. To interpret the moraine ages, we excluded  $^{10}\text{Be}$  ages based on their contribution to the reduced chi-squared ( $\chi_R^2$ ) value (e.g. Kaplan and Miller, 2003). Similar to the approach used by Putnam et al. (2012), we rejected samples until the  $\chi_R^2$  value of  $^{10}\text{Be}$  ages of a moraine was close to one. We also used the stratigraphic order of moraines to reject or retain samples. Fig. DR10 is a comparison of moraine ages based on all  $^{10}\text{Be}$  ages and those based on  $^{10}\text{Be}$  ages were not rejected based on the above criteria. Table DR1 provides data for calculating  $^{10}\text{Be}$  ages. Table DR2 shows the interpreted moraine ages. All ages reported herein are in years before CE 2009 (yr).

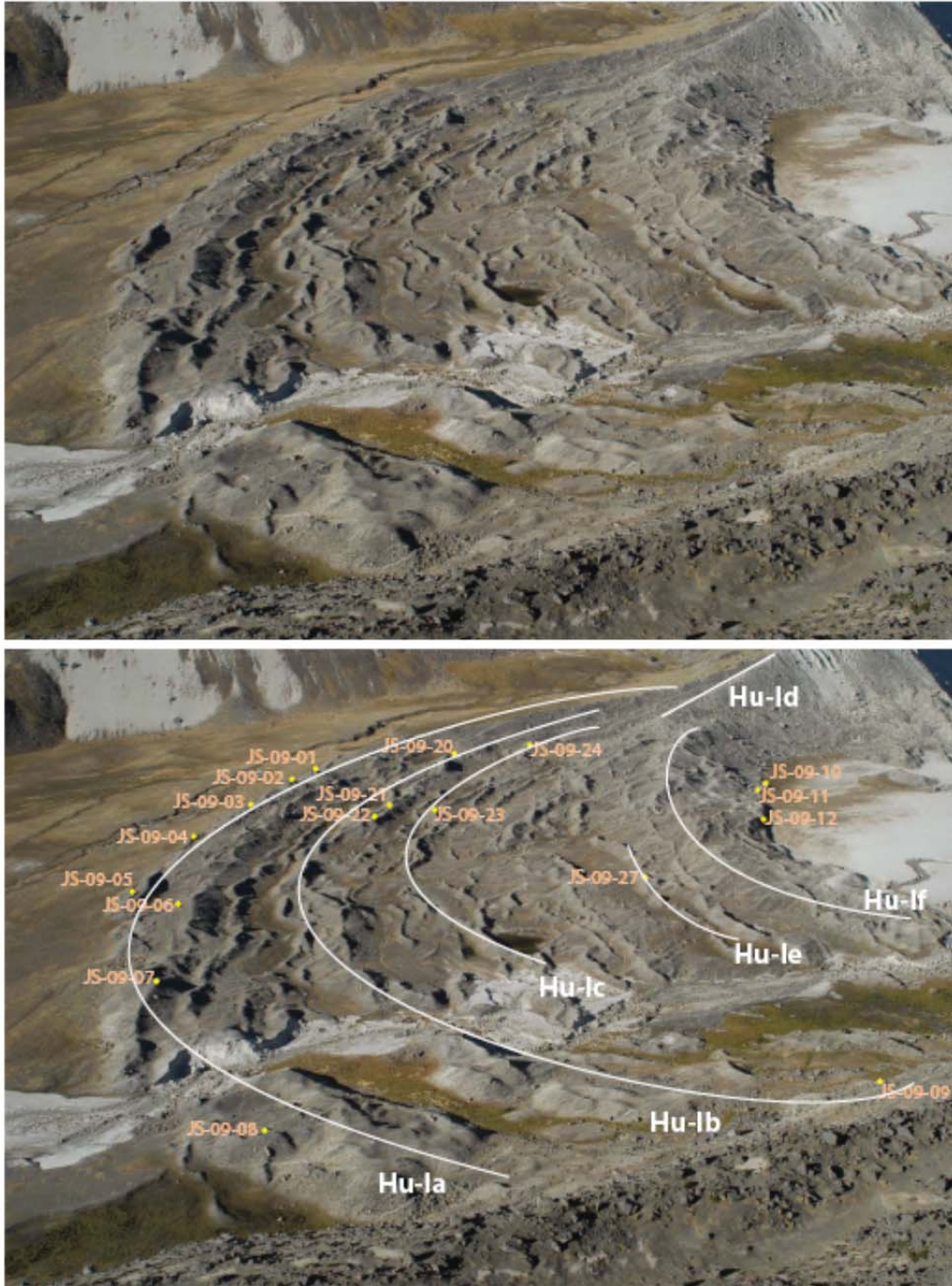


Fig. DR1. Top: Photo of Qori Kalis moraines looking north. Qori Kalis glacier is off the photo to the right (east). Bottom: Approximate moraine positions (labeled Hu-Ia to Hu-If) are marked in white and correspond to the geomorphic map (Fig. 1).  $^{10}\text{Be}$  sample locations are shown in yellow. Crosscutting of the Hu-Ia, Hu-Ib and Hu-Ic moraines by the Hu-Id moraine is apparent in the upper right of the image, indicating that the Hu-Id moraine was deposited by a readvance.

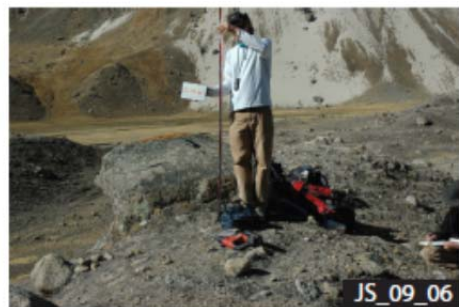
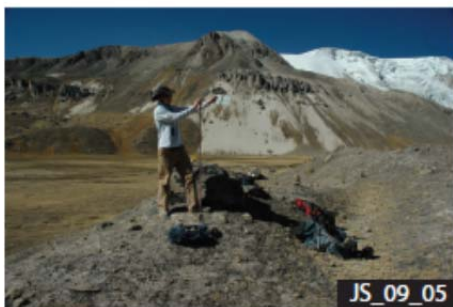
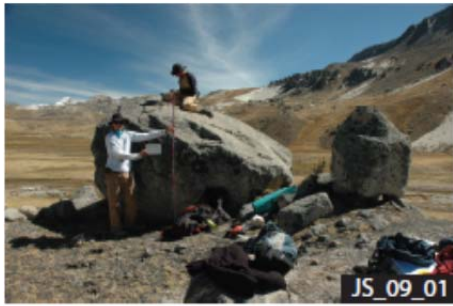


Fig. DR2. Photos of boulders sampled on the Hu-Ia moraine.



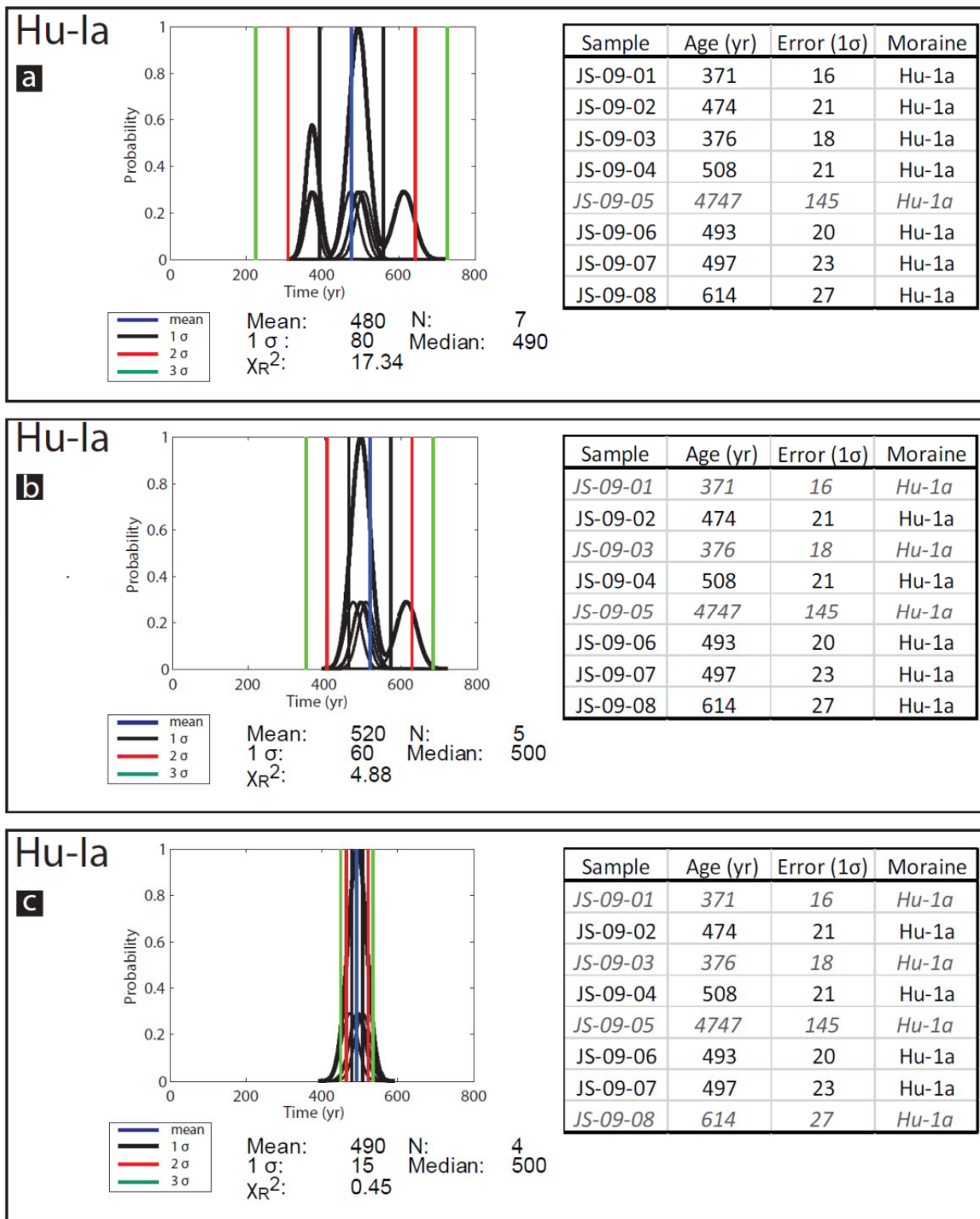


Fig. DR3. Statistics summary for  $^{10}\text{Be}$  ages of the Hu-1a moraine. Plots (a, b and c) are probability distribution functions of  $^{10}\text{Be}$  ages. Vertical lines indicate the means (blue) and standard deviations at the  $1\sigma$ ,  $2\sigma$  and  $3\sigma$  confidence levels (black, red, and green lines, respectively). Statistics associated with the

age populations are shown below the probability distribution functions. Tables (a, b and c) show sample numbers and unrounded  $^{10}\text{Be}$  ages. The samples excluded from the associated statistical analysis are shown in gray italics. a) All  $^{10}\text{Be}$  ages of the Hu-Ia moraine excluding outlier JS-09-05 are included in the calculated statistics. Sample JS-09-05 is a clear outlier and is omitted from moraine age interpretations. With the exclusion of JS-09-05,  $^{10}\text{Be}$  ages ( $n=7$ ) of the Hu-Ia moraine yielded a  $\chi^2_R$  value of 17.34. b) To lower the  $\chi^2_R$  value, we rejected samples JS-09-01 and -03. This yielded a  $\chi^2_R$  value of 4.88. c) Rejecting sample JS-09-08 further lowered the  $\chi^2_R$  value to 0.45; however, there is no geologic evidence that this sample should be removed. Therefore, we retained sample JS-09-08 and consider  $520 \pm 60$  yr as the age of the Hu-Ia moraine.  $^{10}\text{Be}$  ages of boulders on moraines may be apparently older than the age of a moraine if boulders contain  $^{10}\text{Be}$  from a prior period of exposure.  $^{10}\text{Be}$  ages of boulders on moraines may be apparently younger than the age of a moraine if boulders were covered by sediment, snow or vegetation, if they were exhumed from within the landform subsequent to the time of landform deposition, or if sub-aerial erosion acted on the boulder surface. As noted in Table DR1, it is unlikely that boulders sampled for this study were influenced by snow or vegetation cover; however, boulders may have been covered with sediment that was removed prior to our investigation. We were careful to sample only boulders on moraine crests and those that did not show evidence of exhumation.

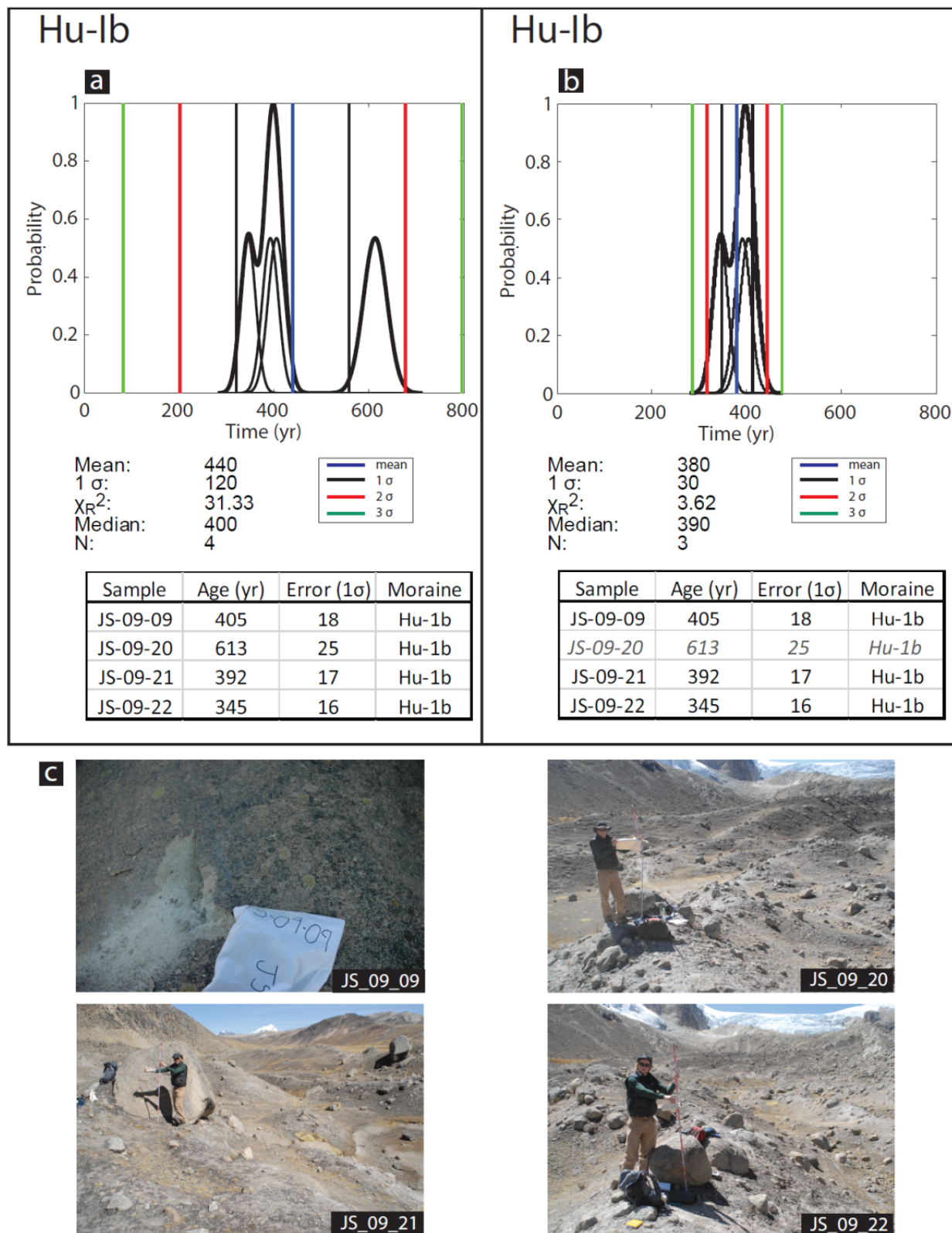


Fig. DR4. Statistics summary for  $^{10}\text{Be}$  ages of the Hu-1b moraine (a and b) and photos of boulders sampled on the Hu-1b moraine (c). a) and b) Same statistical calculations as Fig. DR3. a) All  $^{10}\text{Be}$  ages of

the Hu-Ib moraine are included in the calculated statistics. b) Sample JS-09-20 was excluded because it is older than  $^{10}\text{Be}$  age of the Hu-Ia moraines located downvalley and likely indicates  $^{10}\text{Be}$  from a prior period of exposure (Figs. 1, DR3). Excluding this sample yielded a  $\chi_R^2$  value of 3.62 and an interpreted moraine age of  $380 \pm 30$  yr.

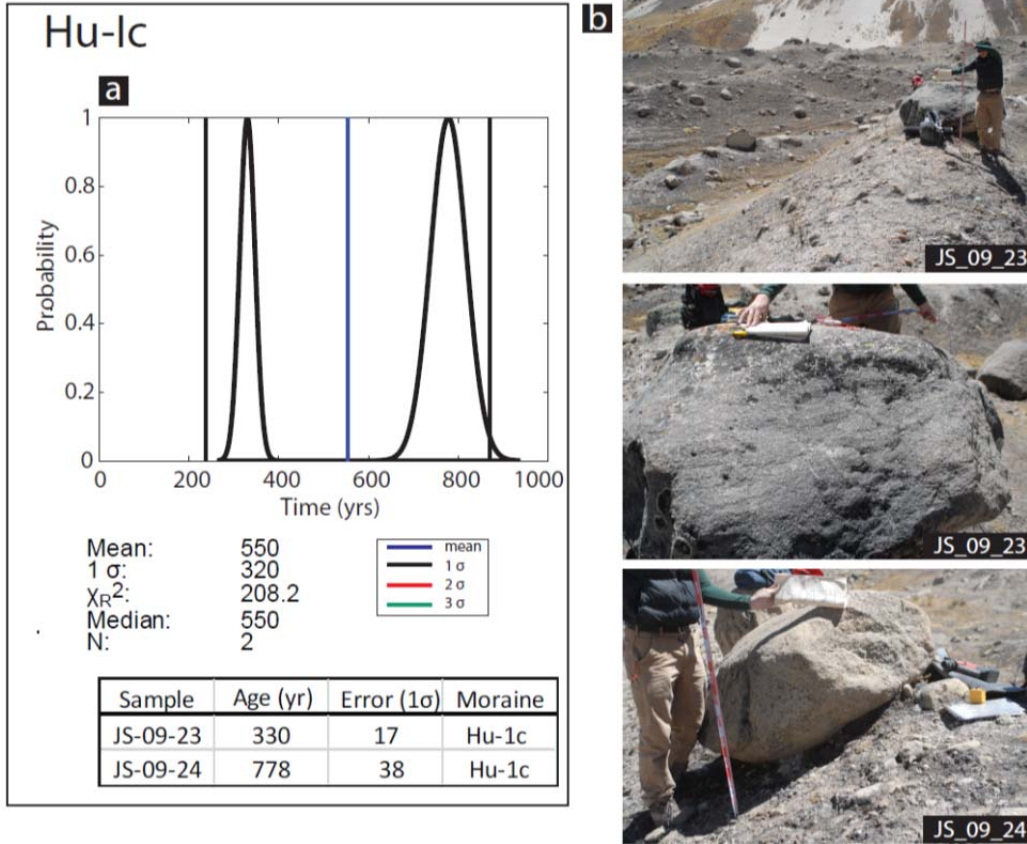


Fig. DR5. Statistics summary for  $^{10}\text{Be}$  ages of the Hu-Ic moraine (a) and photos of boulders sampled on the Hu-Ic moraine (b). a) Same statistical calculations as Fig. DR3. All  $^{10}\text{Be}$  ages of the Hu-Ic moraine are included in the calculated statistics. Sample JS-09-24 was rejected because it is older than  $^{10}\text{Be}$  ages of moraines located downvalley and likely indicates  $^{10}\text{Be}$  from a prior period of exposure (Fig. 1, Figs. DR3, DR4). Since there is only one  $^{10}\text{Be}$  age remaining on this moraine, no  $\chi_R^2$  value is reported. We interpret the age of the Hu-Ic moraine as  $330 \pm 20$  yr. b) The center image (sample JS-09-23) shows a striated boulder surface.

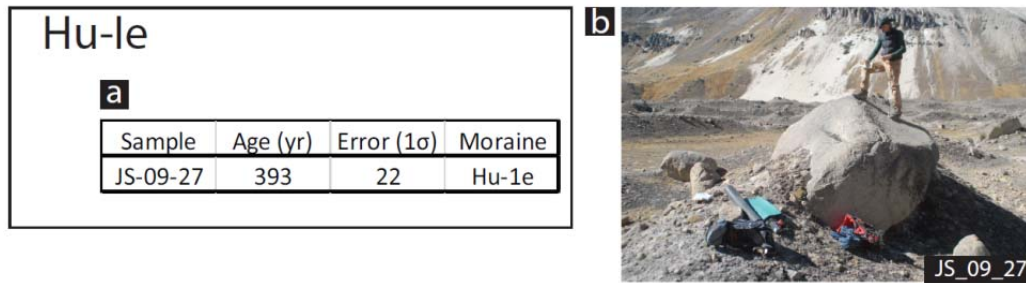


Fig. DR6. Table of the  $^{10}\text{Be}$  age of the Hu-le moraine (a) and photo of the boulder sampled on the Hu-le moraine (b). This sample was rejected because it is out of stratigraphic order with the  $^{10}\text{Be}$  ages of the Hu-Ic and Hu-If moraines. Therefore, we do not interpret an age of the Hu-le moraine.



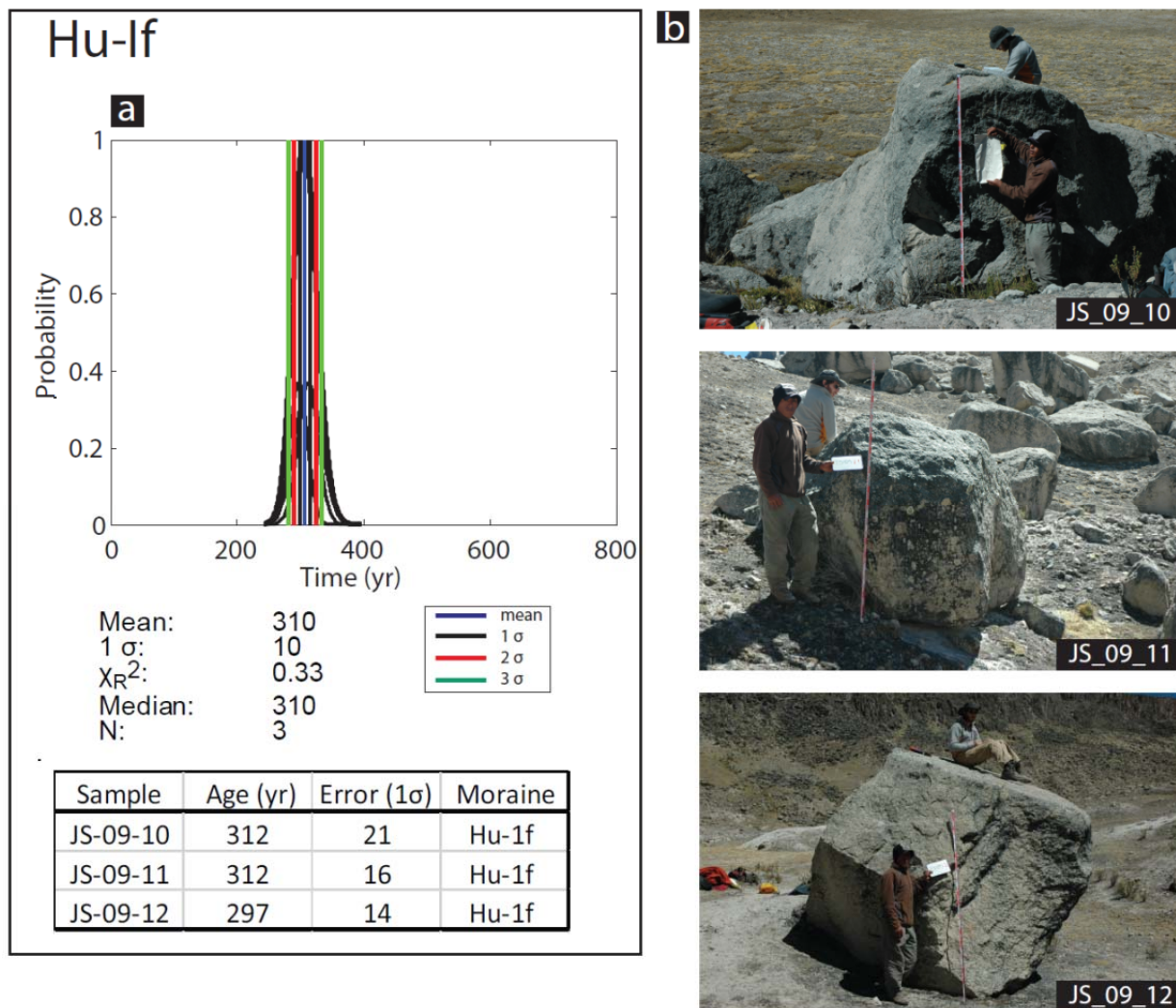


Fig. DR7. Statistics summary for  $^{10}\text{Be}$  ages of the Hu-If moraine (a) and photos of boulders sampled on the Hu-If moraine (b). a) Same statistical calculations as Fig. DR3. We interpret the age of the Hu-If moraine age as  $310 \pm 10$  yr.

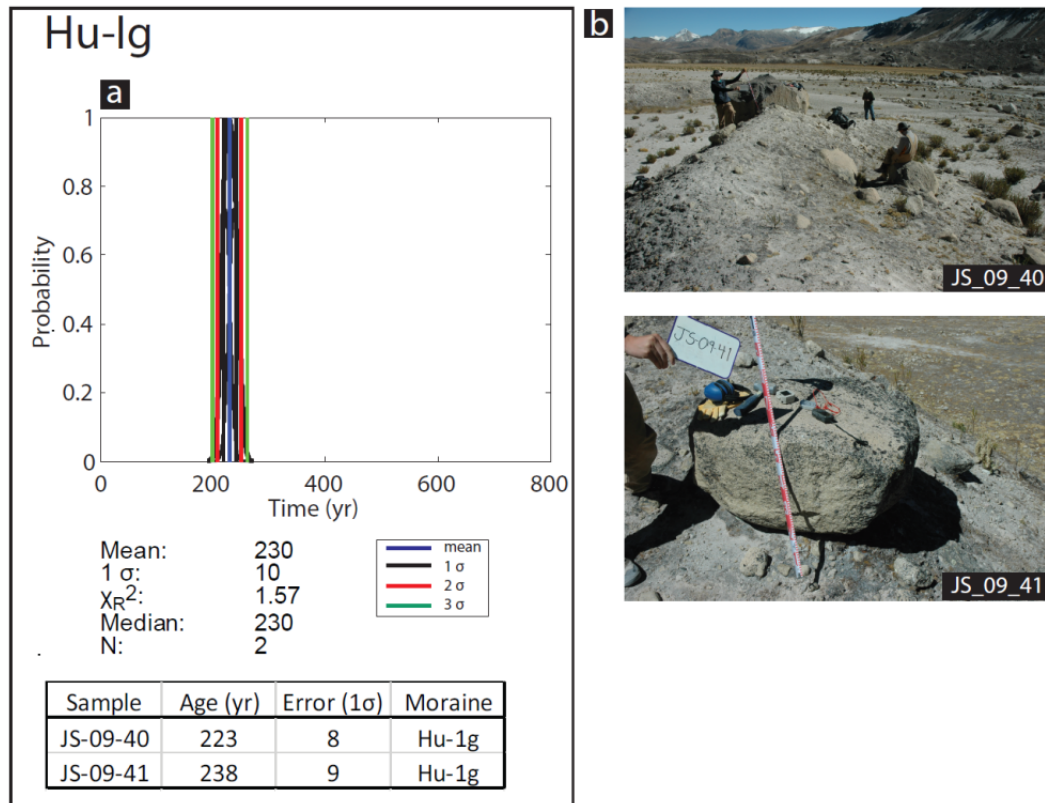


Fig. DR8. Statistics summary for  $^{10}\text{Be}$  ages of the Hu-Ig moraine (a) and photos of boulders sampled on the Hu-Ig moraine (b). a) Same statistical calculations as Fig. DR3. We interpret the age of the Hu-Ig moraine age as  $230 \pm 10$  yr.

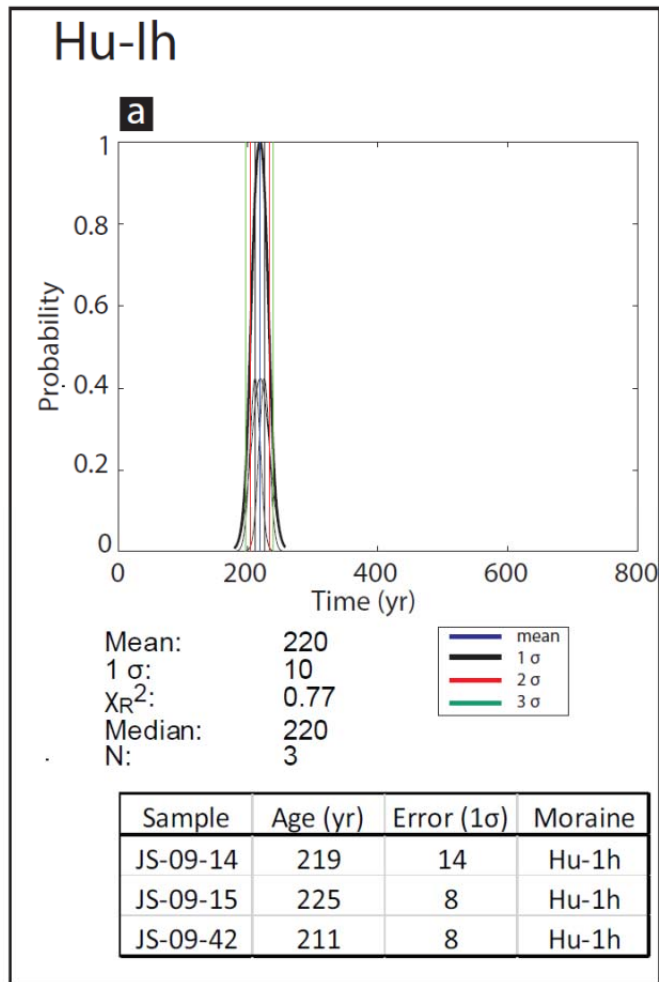


Fig. DR9. Statistics summary for  $^{10}\text{Be}$  ages on the Hu-1h moraine (a) and photos of boulders sampled on the Hu-1h moraine (b). a) Same statistical calculations as Fig. DR3. The Hu-1h moraine marks an ice-margin position similar to that observed in the CE 1963 air photos. The three ages are consistent with one another with a mean age of  $220 \pm 10$  yr. These data likely indicate the ice margin position at CE 1963 and at 220 yr was similar. The  $^{10}\text{Be}$  ages on the Hu-1h moraine are consistent with ice retreat from the Hu-Ig moraines. The ice position between  $\sim 220 \pm 10$  yr and 1963 is unknown but it was likely at or smaller than the Hu-g and Hu-1h moraine positions.

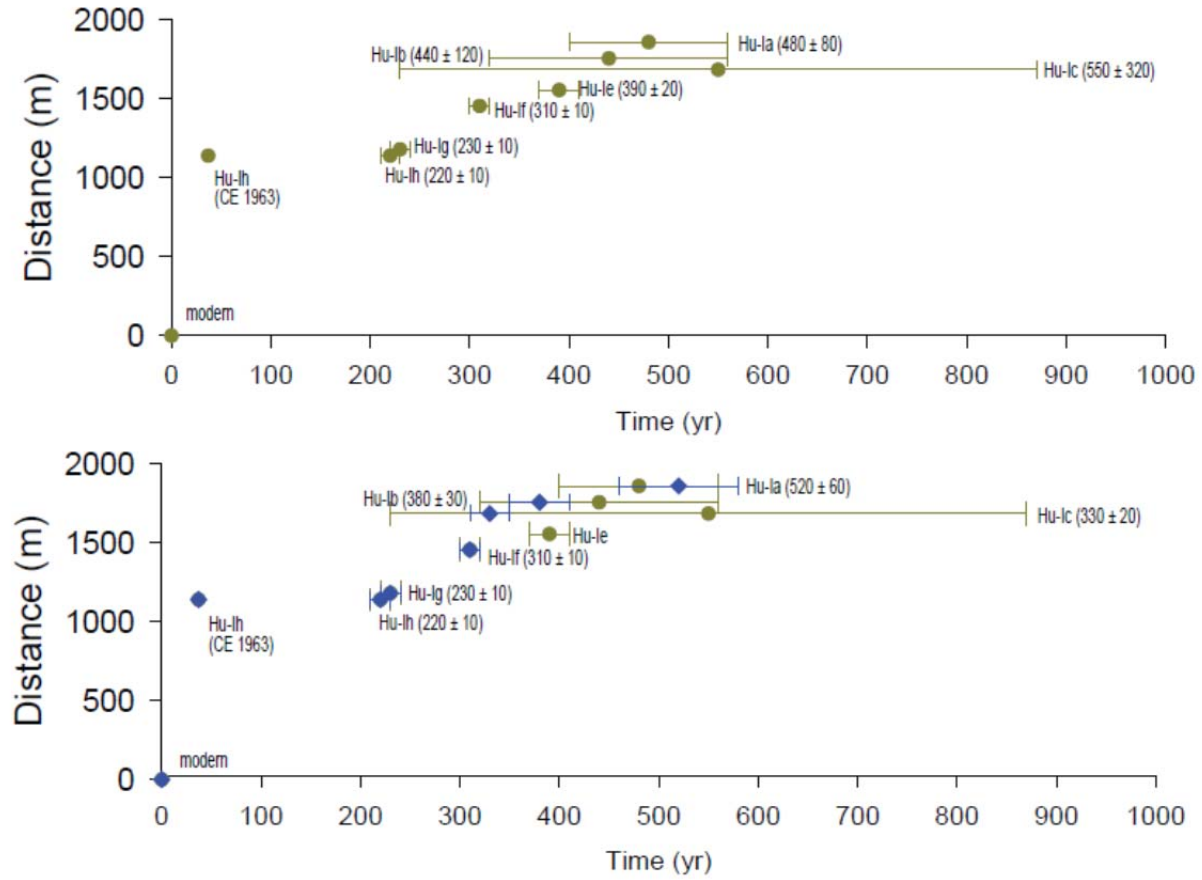


Fig. DR10. Top: The green circles and bars show the mean ages and  $1\sigma$  uncertainties of all  $^{10}\text{Be}$  ages on the Qori Kalis moraines plotted on a graph of distance (down valley from the modern ice margin) and time (excluding sample JS-09-05 on the Hu-la moraine). Bottom: The green bars and circles are the same as the top panel. The blue bars and diamonds show the interpreted mean ages and  $1\sigma$  of moraines based on our exclusion of  $^{10}\text{Be}$  ages. Each moraine is labeled with an interpreted moraine age (Figs. 1, 2 and Table DR2). Both plots indicate younger moraine ages trending upvalley, a pattern consistent with the stratigraphic order of moraines (Figs. 1 and DR1).



TABLE DR1. BOULDER SAMPLE INFORMATION AND  $^{10}\text{Be}$  DATA

Sample ID	CAMS number	Latitude (°N)	Longitude (°W)	Elevation (m asl)	Thickness (cm)	Shielding correction	Quartz wt. (g)	$^9\text{Be}$ Carrier (mg)*	$^{10}\text{Be}/^9\text{Be} \pm 1\sigma$ ( $10^{-14}$ ) <sup>†</sup>	$^{10}\text{Be}$ Conc. $\pm 1\sigma$ (at g <sup>-1</sup> ) <sup>§</sup>	Age $\pm$ Uncertainty (yr) <sup>#</sup>
<b>Hu-Ia</b>											
JS-09-01**	BE30981	-13.9001	-70.8478	4946	2.194	0.998	15.7239	0.2044 <sup>††</sup>	1.8851 $\pm$ 0.0702	16376 $\pm$ 610	370 $\pm$ 20
JS-09-02	BE30982	-13.9000	-70.8481	4947	2.824	0.976	15.9589	0.2052 <sup>††</sup>	2.3666 $\pm$ 0.0886	20338 $\pm$ 761	470 $\pm$ 20
JS-09-03**	BE32652	-13.9000	-70.8487	4944	2.413	0.979	20.6724	0.1985	2.5296 $\pm$ 0.1023	16230 $\pm$ 656	380 $\pm$ 20
JS-09-04	BE30988	-13.9000	-70.8492	4932	2.215	0.929	15.5140	0.2070 <sup>††</sup>	2.3269 $\pm$ 0.0794	20750 $\pm$ 708	510 $\pm$ 20
JS-09-05**	BE32893	-13.9004	-70.8503	4935	1.551	0.902	18.3590	0.1903	27.3528 $\pm$ 0.5147	189421 $\pm$ 3564	4750 $\pm$ 150
JS-09-06	BE32653	-13.9008	-70.8503	4937	1.962	0.928	20.6103	0.1946	3.2032 $\pm$ 0.1037	20207 $\pm$ 654	490 $\pm$ 20
JS-09-07	BE30983	-13.9015	-70.8513	4932	2.318	0.966	15.7745	0.2062 <sup>††</sup>	2.4120 $\pm$ 0.0940	21074 $\pm$ 821	500 $\pm$ 20
JS-09-08	BE32654	-13.9024	-70.8520	4931	2.560	0.877	20.5978	0.1960	3.7128 $\pm$ 0.1390	23613 $\pm$ 884	610 $\pm$ 30
<b>Hu-Ib</b>											
JS-09-09	BE32894	-13.9043	-70.8510	4947	2.173	0.982	14.7193	0.1953	1.9812 $\pm$ 0.0751	17566 $\pm$ 666	410 $\pm$ 20
JS-09-20**	BE32655	-13.9007	-70.8470	4949	2.413	0.908	15.9614	0.1961	2.9962 $\pm$ 0.0973	24599 $\pm$ 799	610 $\pm$ 30
JS-09-21	BE30987	-13.9008	-70.8481	4941	2.507	0.989	15.4057	0.2013 <sup>††</sup>	1.9536 $\pm$ 0.0691	17057 $\pm$ 603	390 $\pm$ 20
JS-09-22	BE32895	-13.9009	-70.8484	4935	2.464	0.902	18.6829	0.1939	1.9692 $\pm$ 0.0797	13659 $\pm$ 552	350 $\pm$ 20
<b>Hu-Ic</b>											
JS-09-23	BE32896	-13.9012	-70.8483	4939	2.180	0.962	17.4106	0.1940	1.8804 $\pm$ 0.0880	14003 $\pm$ 656	330 $\pm$ 20
JS-09-24**	BE32656	-13.9011	-70.8465	4948	1.915	0.920	20.6218	0.1949	5.0284 $\pm$ 0.2136	31766 $\pm$ 1349	780 $\pm$ 40
<b>Hu-Ie</b>											
JS-09-27**	BE32897	-13.9028	-70.8491	4940	2.473	0.988	20.3720	0.1946	2.6716 $\pm$ 0.1378	17057 $\pm$ 880	390 $\pm$ 20
<b>Hu-If</b>											
JS-09-10	BE30984	-13.9029	-70.8467	4938	2.637	0.980	15.8135	0.2056 <sup>††</sup>	1.5434 $\pm$ 0.0988	13412 $\pm$ 858	310 $\pm$ 20
JS-09-11	BE30985	-13.9029	-70.8469	4937	1.973	0.981	15.9237	0.2062 <sup>††</sup>	1.5576 $\pm$ 0.0709	13482 $\pm$ 614	310 $\pm$ 20
JS-09-12	BE30992	-13.9030	-70.8475	4934	2.176	0.994	15.5024	0.2026 <sup>††</sup>	1.4874 $\pm$ 0.0579	12991 $\pm$ 506	300 $\pm$ 10
<b>Hu-Ig</b>											
JS-09-40	BE32898	-13.9051	-70.8465	4944	2.718	0.975	52.8348	0.1986	3.7995 $\pm$ 0.1036	9546 $\pm$ 260	220 $\pm$ 10
JS-09-41	BE32899	-13.9056	-70.8470	4935	3.656	0.911	47.8400	0.1943	3.4839 $\pm$ 0.1049	9454 $\pm$ 285	240 $\pm$ 10
<b>H-Ih</b>											
JS-09-42	BE32659	-13.9040	-70.8445	4942	3.117	0.950	68.2451	0.1957	4.5794 $\pm$ 0.1445	8774 $\pm$ 277	210 $\pm$ 10
JS-09-14	BE32657	-13.9053	-70.8459	4948	3.458	0.983	71.8010	0.1953	5.1839 $\pm$ 0.3026	9425 $\pm$ 550	220 $\pm$ 10
JS-09-15	BE32658	-13.9057	-70.8455	4939	2.421	0.978	71.6553	0.1946	5.3286 $\pm$ 0.1413	9670 $\pm$ 256	230 $\pm$ 10

\* All samples were prepared at Dartmouth College. Samples marked with an “††” were prepared using the beryllium carrier “Lamont Carrier 5.1”. All other samples were prepared using the beryllium carrier “4G Dartmouth Carrier.”

<sup>†</sup> Beryllium measurements were made relative to the 07KNSTD3110 standard (Nishiizumi et al., 2007). Internal AMS uncertainties are reported.

<sup>§</sup>  $^{10}\text{Be}$  ages assume zero erosion and are calculated using the CRONUS-Earth online calculator (Balco et al., 2008) with a locally calibrated production rate ( $P_{\text{Quel}}$ ;  $3.78 \pm 0.09$  [St] atoms g<sup>-1</sup> yr<sup>-1</sup>; Kelly et al. (2013).

<sup>#</sup> External uncertainties for  $^{10}\text{Be}$  age calculations, rounded to the nearest ten years.

\*\* Sample excluded from interpreted moraine ages (DR3-10).

Table DR1.  $^{10}\text{Be}$  sample data and calculated  $^{10}\text{Be}$  surface exposure ages. Shown are Dartmouth sample numbers; Center for Accelerator Mass Spectrometry (CAMS) sample numbers; sample latitudes, longitudes and elevations; sample thicknesses; correction factors for sample surface slopes and topographic shielding (Shielding correction); sample quartz amounts (Quartz wt.);  $^9\text{Be}$  carrier amounts; measured  $^{10}\text{Be}/^9\text{Be}$  ratios and  $1\sigma$  uncertainties; calculated  $^{10}\text{Be}$  concentrations and  $1\sigma$  uncertainties ( $^{10}\text{Be}$  conc.); and calculated  $^{10}\text{Be}$  ages and external uncertainties (Age).

We collected ~1 kg samples of rock from the center of flat or gently sloping boulder surfaces using a hammer and chisel or the drill-and-blast method of Kelly (2003). Before sample extraction, we used a compass and clinometer to measure the strike and dip of the sample surface as well as horizon angles at 20° azimuth intervals. We used these measurements to calculate topographic shielding. There was no evidence for post-depositional boulder-surface erosion. On some boulder surfaces we observed glacial polish. We considered the effects of snow, sediment and vegetation cover of boulders to be negligible. Due to intense solar radiation, snow cover does not persist on the landscape. We did not observe sediment cover on any boulders. There was little-to-no vegetation on the moraines. All samples were prepared in the cosmogenic nuclide laboratory at Dartmouth College following methods modified from Stone et al. (2001) and similar to those used by Schafer et al. (2009).  $^{10}\text{Be}/^9\text{Be}$  ratios were measured at the CAMS at Lawrence Livermore National Laboratory (LLNL) relative to the 07KNSTD3110 standard (Nishiizumi et al., 2007).

**TABLE DR2. INTERPRETED MORaine AGES**

Moraine	Mean $\pm$ 1 $\sigma$ (yr)
Hu-Ia	520 $\pm$ 60
Hu-Ib	380 $\pm$ 30
Hu-Ic	330 $\pm$ 20
Hu-Id*	350-300
Hu-If	310 $\pm$ 10
Hu-Ig	230 $\pm$ 10
Hu-Ih	220 $\pm$ 10 and CE1963

\*See figure caption for explanation of age assignment.

Table DR2. Summary of interpreted moraine ages. The Hu-Id moraine is not directly dated but it crosscuts the Hu-Ia, Ib and Ic moraines (Figs.1 and DR1) and must have been deposited prior to the deposition of the Hu-Ie moraine. The closest minimum-limiting age bracket for the Hu-Id moraine is provided by the age of the Hu-If moraine. Using the ages from of the Hu-Ic (330 $\pm$ 20 yr) and the Hu-If (310 $\pm$ 10 yr) moraines, we assigned the Hu-Id moraine an age of 350-300 yr.

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