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The Ries impact, a double layer rampart crater on Earth.

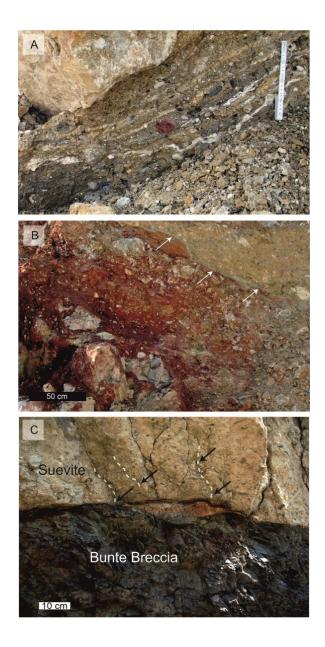
2 Supplementary material

3 Macro- and microscopic analyses of Bunte Breccia samples

Intensive macro - and microscopic analyses of several Bunte Breccia samples of two drilling 4 5 sites (Otting and Itzing) and of surface outcrops revealed that a large amount of water was 6 entrained in the Bunte breccias during flow (Sup. Fig. s1, s2). Evidence for fluidized 7 behaviour comes from blocks that are coated with rims of clays in which smaller clastic debris exhibit alignments. This may indicate the presence of water saturated clays that accreted 8 9 around the blocks during gliding and rolling transport. Evidence also comes from very thin 10 clays with very high aspect ratios that were ductily rolled out. Another indicator is that the vertical vent pipes that occur within the overlying Suevite originate directly from the contact 11 to the Bunte breccias. These pipes likely vented steam from water saturated Bunte Breccia 12 13 that was covered with hot Suevite (Wittmann and Kenkmann, 2007). Boyce et al. (2012) 14 proposed that martian pits, found in thin ejecta material outside the crater, on terraced blocks 15 in the interior of the crater and on ejecta blankets near the crater rim, were formed due to rapid degassing of water from the pitted material. It is supposed that this process is similar to 16 17 the degassing of the fall fall-suevite of the Ries impact structure (Boyce et at., 2012, 18 Tornabene et al., 2012).

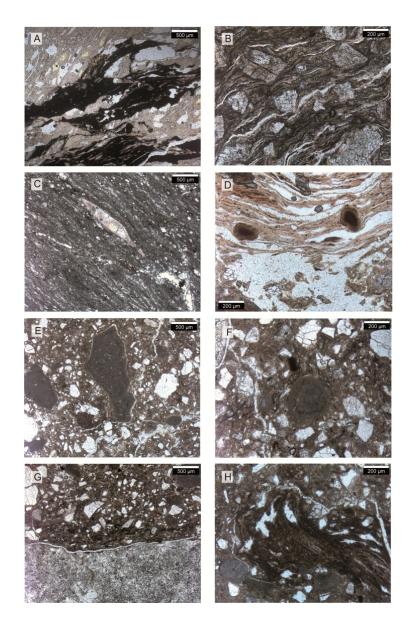
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Supplementary Figure s1: Macroscopic fabrics within Bunte Breccia deposits of the continuous ejecta blanket. (a) Gundelsheim, active quarry (GM Gundelsheimer Marmorwerk), 21 km E of crater centre: Variegated clasts are embedded in a clay-rich matrix that displays intensive fluided-assisted ductile flow. The matrix flow led to compositional banding and small-scale disharmonic folding of the matrix. Clasts were mechanically abraded. (b) MaurenBräulesberg, active quarry (Märker Zement), 15 km SE of crater centre: two types of matrixsupported breccias are in sharp contact (arrows) to each other. The contact zone is a strongly

localized shear plane that developed during the final movements of the Bunte Breccia (c)
Aumühle, active quarry (Märker Zement), 9 km NE of crater centre: Sharp contact plane
between the Bunte breccia deposits, here represented by dark claystone of Liassic age, and
Suevite. Degassing pipes are emanating from the contact plane and penetrate into the hanging
wall Suevite deposits (dashed lines). The pipes indicate the presence of water in the Bunte
Breccia and subsequent vaporization upon deposition of the hot Suevite.





40 Supplementary Figure s2: Microscopic fabrics within the Bunte Breccia. a-c) drill core
41 Itzing, 20.8 km of the crater centre, d-h) drill core Otting, 17.4 km of the crater centre. a) A

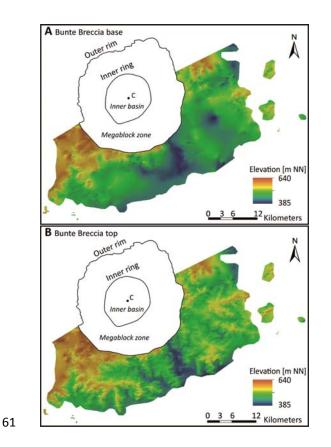
42	ductile flowing limestone-clay-iron-hydroxide assemblage is crosscut by a localized shear
43	plane. B) Clasts of limestone are flowing in a clay-dominated matrix. C) Limestone
44	porphyroclasts are embedded in a fine-grained clay matrix. Porphyroclasts can be used as
45	shear sense indicators (σ -clasts) and suggest simple shear and a laminar low viscosity ductile
46	flow that must have been assisted by a fluid phase. D) Rounded clasts in a ductile lime-clay
47	matrix show a core and mantle structure. E+F) Typical assemblage of sub-angular
48	sedimentary target clasts suspended in Bunte Breccia matrix. The coated clasts in the middle
49	indicate that clays are mantling particles during a rolling process. G) Preferred alignment of
50	particles and flow lines near the contact to a larger limestone clast. H) Breccia in breccias
51	pattern with a chaotic assemblage of ductile deformed and folded clasts in a clay-rich matrix.

52	ASTER Global	Digital Elevation	Model – Validation	results of GDME 2
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			Version 2 - GDME2
Horizontal Error			0.13 arc-sec. to west
			0.19 arc-sec. to north
	Flat and open area	offset	-0.7 m
	(rice farm)	SD	5.9 m
Elevation		RMSE	6.1 m
Error	Mountainous area largely	offset	+7.4 m
	covered by forest	SD	12.7 m
		RMSE	15.1 m
Horizontal Resolution			2.4 arc-sec (72m)

Supplementary Table s1: Validation results from a Japanese study of the Global Digital
Elevation Model 2 (GDEM2) ASTER dataset that was used for our interpolation study of the
ejected Bunte Breccia material (one arc-second corresponds to 30 meters) (modified after
Tachikawa et al., 2011).

60 Interpolation results



62 Supplementary Figure s3: Results of the interpolated morphology of the (a) Bunte Breccia

63 base and (b) Bunte Breccia upper surface.

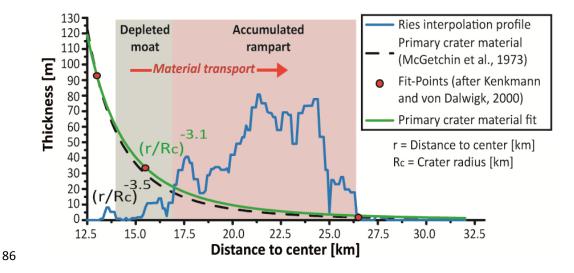
71 Ejecta layer morphometric parameters

	Average ejecta mobility / crater radius	Average rampart width [km]	Average rampart width / crater radius
Mars SLE	1.27	1.04	0.17
Mars DLE inner layer	1.37	5.28	0.66
Mars DLE outer layer	3.17	1.29	0.17
Mars MLE inner layer	1.25	0.98	0.11
Mars MLE outer layer	2.31	1.13	0.12
GRLE SLE (Nergal)	1.00	0.80	0.20
GRLE DLE inner layer	0.86	8.24	0.43
GRLE DLE outer layer	1.86	4.14	0.12
Earth $(Lonar)^{*1}$	2.34	0.30	0.31
Earth (Ries) inner layer ^{*2}	1.12	8.70	0.67
Earth (Ries) outer layer ^{*2}	2.36	?	?

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Supplementary Table s2: Ejecta layer morphometric parameters for fluidized ejecta craters
with ramparts found on Mars, Ganymede (GRLE) and Earth (SLE = single-layer ejecta, DLE
= double-layer ejecta, MLE = multiple-layer ejecta) (modified after Boyce et al., 2010;
*¹Maloof et al., 2009; *²this study).

85 Material difference calculation



Supplementary Figure s4: Minimum ejecta thickness of one interpolated Bunte Breccia
thickness profile (blue line, profile orientation is shown in Fig. 1) of the Ries impact crater
compared to the thickness variation of ejected primary crater material (dashed line, calculated
after McGetchin et al., 1973) and recalculated and fitted to three fit-points (red dots) after
Kenkmann and von Dalwigk (2000) (green line).

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