## GSA DATA REPOSITORY 2011271

Weitz et al.

## Supplementary materials for "Diverse mineralogies in two troughs of Noctis Labyrinthus, Mars"

## **CRISM Data Processing**

We processed CRISM hyperspectral images using standard procedures developed by the team [Murchie et al., 2007; 2009a]. The CRISM data were converted to I/F by subtracting the instrument background, and dividing by an internal calibration standard and solar irradiance. The resulting I/F was divided by the cosine of the incidence angle (from MOLA) in order to correct the illumination and atmospheric contributions were minimized by dividing by an atmospheric transmission spectrum over Olympus Mons. Finally, a cleaning algorithm was applied to the images to remove noise and large spikes within the data that are due to instrument effects [Parente, 2008].

Within any unit,  $\sim 10-100$  pixels were averaged and this result was then divided by an average of spectrally neutral pixels along the same column in the CRISM scene. These CRISM ratioed spectra were then compared to laboratory data of pure minerals for potential matches; however, difficulties associated with identifying a perfect match to laboratory spectra can be attributed to subpixel mixing in the CRISM data. Our observations of diverse morphologies within several units based upon HiRISE images supports this subpixel mixing at the CRISM scale ( $\sim 20$  m/pixel).

Parameter maps corresponding to diagnostic mineralogies were generated to map the distribution of minerals such as clays, sulfates, pyroxenes, and olivines [Pelkey et al., 2007; Murchie et al., 2009b] and aid in the selection of spectra. Spectral parameter color maps (Figures 2A, 4A) were produced as follows: the 2.3  $\mu$ m drop in reflectance, due to Fe/Mg-OH vibrations (D2300), is shown in red; the convexity around 2.3  $\mu$ m in very hydrated phases due to strong H<sub>2</sub>O absorptions at 1.9 and 2.4  $\mu$ m (SINDEX) is shown in green; and the 1.9  $\mu$ m band depth, due to combinations of H<sub>2</sub>O bending and stretching vibrations (BD1900) is shown in blue.



Figure DR1. (a) Portion of HiRISE image ESP\_015975\_1695 showing (b) blowup views of Trough 1 units LTL(Su), LT(K), and LTL3. (c) Rounded blocks of LTL(Su) mixed into material of unknown lithology along the southeastern portion of the trough.



Figure DR2. (a) Portion of HiRISE image ESP\_015975\_1695 showing blowup views (be) of Trough 1 units LT(Db), LTL1, LTL2, MTL1, LTL(Su), LTL(FeSm) and Brecciated.



Figure DR3. (a) Portion of HiRISE image ESP\_016898\_1695 showing blowup views of Trough 1 units. (b) HiRISE stereo anaglyph of same region shown in a. (c) Blowups of LTL4 and (d) LTL(Su), LTL(FeSm), and MT(HdSi). Notice the finger-like extensions of LTL(FeSm) filling topographic lows upslope along unit MT(HdSi).



Figure DR4. (A) CRISM spectra of all units in Trough 1 that display hydration absorptions. (B) Selected laboratory spectra that have similar features and absorptions to the units in Trough 1. The FeMg-smectite sample is from Flagstaff Hill, CA, and the spectrum is from Bishop et al. [2008]. The hydrated silica spectrum is from Bishop et al. [2005a] and the sample is an altered ash near S vents inside Kilaeau and contains a large amount of opal-A. The halloysite spectrum is from Bishop and Murad [1996], the ferricopiapite and szomolnokite spectra are from Lane et al. [2004] and the spectrum of FeOHSO<sub>4</sub> created through dehydration of ferricopiapite is from Bishop et al. [2009].



Figure DR5. (a) Portion of false-color HiRISE image PSP\_005400\_1685 showing blowup views (b-e) of Trough 2 units LTL10, LTL9, LTL8, and LTL(FeSm).



Figure DR6. (a) Portion of HiRISE image PSP\_005400\_1685 showing blowup views (b-e) of Trough 2 units LTL8, LTL7, LTL(Su), LTL(Op+Fe), LTL2, LTL1, and MTL.



Figure DR7. (a) Portion of HiRISE image PSP\_005400\_1685 showing blowup views (bf) of Trough 2 units LTL5, LTL6, LTL(FeSm1), LTL(FeSm2), and LTD.



Figure DR8. (a) Portion of false-color HiRISE image ESP\_014353\_1685 showing blowup views (b,c) of Trough 2 units LTL3 and LTL4.



Figure DR9. (A) CRISM spectra of all units in Trough 2 that display hydration absorptions. (B) Selected laboratory spectra that have similar features and absorptions to the units in Trough 2. The FeMg-smectite spectrum is from Bishop et al. [2008] and the beidellite spectrum is from Bishop et al. [2011]. The SBCa-1 beidellite sample is a Ca-

saturated <0.2  $\mu$ m sediment extraction of a natural sample from CA [Gates, 2005]. The opal spectrum is from McKeown et al. [2011] and is opal-CT from Virgin Valley, NV. The kaolinite KGa-1 is from Washington County, GA and the spectrum is from Bishop et al. [2008b]. The hexahydrite spectrum is from Cloutis et al. [2006], the schwertmannite and jarosite spectra are from Bishop and Murad [1996], the ferricopiapite spectrum is from Lane et al. [2004] and the rozenite spectrum is from Bishop et al. [2005b].

## REFERENCES

- Bishop, J.L., and Murad, E., 1996, Schwertmannite on Mars? Spectroscopic analyses of schwertmannite, its relationship to other ferric minerals, and its possible presence in the surface material on Mars, *in* Dyar, M.D., McCammon, C., and Schaefer, M.W., eds., Mineral Spectroscopy: A tribute to Roger G. Burns, Volume Special Publication Number 5: Houston, TX, The Geochemical Society, p. 337-358.
- Bishop, J.L., and Murad, E., 2005, The visible and infrared spectral properties of jarosite and alunite: American Mineralogist, v. 90, p. 1100-1107.
- Bishop, J.L., Schiffman, P., Lane, M.D., and Dyar, M.D., 2005a, Solfataric alteration in Hawaii as a mechanism for formation of the sulfates observed on Mars by OMEGA and the MER instruments, Lunar Planet. Sci. XXXVI.: Lunar Planet. Inst., Houston, p. Abstract #1456.
- Bishop, J.L., Dyar, M.D., Lane, M.D., and Banfield, J.F., 2005b, Spectral identification of hydrated sulfates on Mars and comparison with acidic environments on Earth: International Journal of Astrobiology, v. 3, p. 275-285.
- Bishop, J.L., Lane, M.D., Dyar, M.D., and Brown, A.J., 2008a, Reflectance and emission spectroscopy study of four groups of phyllosilicates: Smectites, kaolinite-serpentines, chlorites and micas Clay Minerals, v. 43, p. 35-54.
- Bishop, J.L., Noe Dobrea, E.Z., McKeown, N.K., Parente, M., Ehlmann, B.L., Michalski, J.R., Milliken, R.E., Poulet, F., Swayze, G.A., Mustard, J.F., Murchie, S.L., and Bibring, J.-., P., 2008b, Phyllosilicate diversity and past aqueous activity revealed at Mawrth Vallis, Mars: Science, v. 321, p. DOI: 10.1126/science.1159699, pp. 830-833.
- Bishop, J.L., Parente, M., Weitz, C.M., Noe Dobrea, E.Z., Roach, L.A., Murchie, S.L., McGuire, P.C., McKeown, N.K., Rossi, C.M., Brown, A.J., Calvin, W.M., Milliken, R.E., and Mustard, J.F., 2009, Mineralogy of Juventae Chasma: Sulfates in the Light-toned Mounds, Mafic Minerals in the Bedrock, and Hydrated Silica and Hydroxylated Ferric Sulfate on the Plateau: Journal of Geophysical Research, v. 114, p. doi:10.1029/2009JE003352.
- Bishop, J.L., Gates, W.P., Makarewicz, H.D., McKeown, N.K., and Hiroi, T., 2011, Reflectance spectroscopy of beidellites and their importance for Mars: Clays and Clay Minerals, p. under revision.
- Clark, R.N., Swayze, G.A., King, T.V.V., Gallagher, A., and Calvin, W.M., 1993, Digital Spectral Library: Version 1: 0.2 to 3.0-μm.: USGS Open File Report 93-592.
- Cloutis, E.A., Hawthorne, F.C., Mertzman, S.A., Krenn, K., Craig, M.A., Marcino, D., Methot, M., Strong, J., Mustard, J.F., Blaney, D.L., Bell, J.F., III, and Vilas, F., 2006, Detection and discrimination of sulfate minerals using reflectance spectroscopy: Icarus, v. 184, p. 121-157.

- Gates, W.P., 2005, Infrared spectroscopy and the chemistry of dioctahedral smectites, *in* Kloprogge, J.T., ed., The Application of Vibrational Spectroscopy to Clay Minerals and Layered Double Hydroxides, Volume 13: Aurora, CO, Clay Minerals Society, p. 125-168.
- Lane, M.D., Dyar, M.D., and Bishop, J.L., 2004, Spectroscopic evidence for hydrous iron sulfate in the Martian soil: Geophysical Research Letters, v. 31, p. L19702, doi:10.1029/2004GL021231.
- McKeown, N.K., Bishop, J.L., Cuadros, J., Hillier, S., Amador, E., Makarewicz, H.D., Parente, M., and Silver, E., 2011, Interpretation of reflectance spectra of mixtures of clay minerals and silica: implications for Martian clay mineralogy at Mawrth Vallis: Clays and Clay Mineral, in review.
- Murchie, S., Arvidson, R., Bedini, P., Beisser, K., Bibring, J.-P., Bishop, J., Boldt, J., Cavender, P., Choo, T., Clancy, R.T., Darlington, E.H., Des Marais, D., Espiritu, R., Fort, D., Green, R., Guinness, E., Hayes, J., Hash, C., Heffernan, K., Hemmler, J., Heyler, G., Humm, D., Hutcheson, J., Izenberg, N., Lee, R., Lees, J., Lohr, D., Malaret, E., Martin, T., McGovern, J.A., McGuire, P., Morris, R., Mustard, J., Pelkey, S., Rhodes, E., Robinson, M., Roush, T., Schaefer, E., Seagrave, G., Seelos, F., Silvergate, P., Slavney, S., Smith, M., Shyong, W.-J., Strohbehn, K., Taylor, H., Thompson, P., Tossman, B., Wirzburger, M., and Wolff, M., 2007, Compact Reconnaissance Imaging Spectrometer for Mars (CRISM) on Mars Reconnaissance Orbiter (MRO): Journal of Geophysical Research, v. 112, p. doi:10.1029/2006JE002682.
- Murchie, S. L., et al. 2009a, Compact Reconnaissance Imaging Spectrometer for Mars investigation and data set from the Mars Reconnaissance Orbiter's primary science phase, J. Geophys. Res., 114, E00D07, doi:10.1029/2009JE003344.
- Murchie, S.M., et al., 2009b, A synthesis of martian aqueous mineralogy after 1 Mars year of observations from the Mars Reconnaissance Orbiter: Journal of Geophysical Research, 114, E00D06, doi:10.1029/2009JE003342.
- Parente, M. (2008), A new approach to denoising CRISM images, paper presented at 39<sup>th</sup> Lunar Planet Science Conf, LPI, Houston, TX, abs# 2528.
- Pelkey, S.M., et al., 2007, CRISM multispectral summary products: Parameterizing mineral diversity on Mars from reflectance: Journal of Geophysical Research, v. 112, p. E08S14, doi:10.1029/2006JE002831.