

Table DR1: OSL results,  $D_e$  breakup for Paran samples.

Lab No.	Grain size ( $\mu\text{m}$ )	Aliquot size (mm)	No. of discs	$D_e$ (Gy)
PS-40	74-105	5	9	$1.55 \pm 0.18$
	150-177	5	9	$1.54 \pm 0.44$
	150-177	2	23	$1.38 \pm 0.74$
PS-41	88-125	5	13	$6.1 \pm 1.3$
	88-125	2	22	$4.9 \pm 1.2$
PS-42	74-105	5	9	$18.3 \pm 2.8$
	150-177	5	9	$21.0 \pm 9.1$
	150-177	2	24	$21.9 \pm 10.4$
PS-43	74-105	5	9	$35.9 \pm 11.4$
	150-177	5	9	$37.7 \pm 12.2$
	150-177	2	22	$32.0 \pm 15.8$
PS-44	88-125	5	13	$66.5 \pm 20.7$
	88-125	2	24	$50.2 \pm 23.8$
PS-45	74-105	5	9	$76.8 \pm 16.9$
	150-177	5	9	$83.4 \pm 32.9$
	150-177	2	24	$93.7 \pm 54.3$
PS-46	74-105	5	9	$112 \pm 23$
	150-177	5	9	$120 \pm 34$
PS-48	74-105	5	9	$185 \pm 59$
	150-177	5	9	$215 \pm 58$

$D_e$  values for different grain sizes or aliquot sizes. The 2 mm aliquots were measured using a standardized growth curve based on the average of 5 aliquots.

The averages presented here are unweighted means. In Table 1 the central age model (Galbraith et al. 1999) was used to calculate the most representative  $D_e$  from all measurements.

Table DR2: Luminescence ages for Har Harif samples (from Crouvi et al., 2009)

Sample	Depth (m)	K (%)	U (ppm)	Th (ppm)	Ext. $\alpha$ $\mu\text{Gy/a}$	Ext. $\beta$ $\mu\text{Gy/a}$	Ext. $\gamma$ + Cos. $\mu\text{Gy/a}$	Cos. $\mu\text{Gy/a}$	Total dose $\mu\text{Gy/a}$	Aliquots used	D <sub>e</sub> (Gy)	Age (ka)
HR-1 <sup>(2)</sup>	0.15	1	2.5	6.4	12	1155	951	-	2118±99	27/29	29.1±0.6	<b>13.7±0.7</b>
HR-13 <sup>(1)</sup>	0.27	0.87	1.7	4.7	9	923	606	243	1782±27	11/12	44.6±1.1	<b>25.1±0.7</b>
HR-5	0.40	0.96	0.9	5.6	7	904	1023	-	1934±105	17/18	63.4±1.8	<b>32.8±2.0</b>
HR-6	0.77	0.91	1.7	5.4	9	968	966	-	1943±100	13/13	81.6±4.0	<b>42.0±3.0</b>
HR-2 <sup>(1)</sup>	0.90	1	1.7	5.5	9	1007	659	181	1856±29	18/18	90.2±2.7	<b>48.6±2.7</b>
HR-7	1.65	0.95	1.8	5.4	9	984	903	-	1896±93	13/13	117±4.2	<b>61.5±3.8</b>
HR-16	1.65	0.94	2.1	6.2	11	1033	929	-	1973±96	12/12	127±3.2	<b>64.4±3.6</b>
HR-8	1.90	0.67	1.6	4.4	8	749	760	-	1517±79	8/13	117±5.0	<b>77.4±5.2</b>
HR-9	2.35	0.67	1.7	5.0	8	775	779	-	1563±81	13/13	118±6.7	<b>72.5±5.8</b>
HR-3 <sup>(1)</sup>	2.55	0.74	1.9	5.1	9	849	603	147	1608±27	11/18	147±3.9	<b>91.3±2.9</b>
HR-10	2.75	0.74	1.9	5.1	9	849	790	-	1648±83	13/13	120±8.3	<b>73.0±6.2</b>
HR-4 <sup>(1)</sup>	2.95	0.7	2.1	5.2	10	850	620	135	1615±27	17/18	154±6.6	<b>95.6±4.4</b>
HR-11	3.05	0.78	2.0	5.6	10	900	803	-	1713±84	7/13 6/13	225±10 357±15	<b>131±8.7</b> 208±13
HR-12	3.45	0.57	1.5	3.6	7	651	685	-	1343±73	12/13	242±4.5	<b>181±10</b>
HR-14	3.90	0.50	1.8	3.8	8	646	633	-	1287±67	17/17	233±6.7	<b>181±11</b>
HR-15	4.30	0.46	1.8	3.5	8	613	590	-	1210±63	12/12	395±22	<b>326±25</b>

Horizontal lines separate between the different units. Water contents were estimated at 3±1 % for 0-1 m depth and 5±1 % below that. Gamma + cosmic dose rates were measured in the field using a portable gamma counter, except for samples marked with <sup>(1)</sup>, for which gamma was calculated from the radioelements and the cosmic dose estimated from burial depth. Measurements were carried out on 74-125  $\mu\text{m}$  quartz using a modified SAR protocol (Murray and Wintle, 2000). Preheat temperatures were 10 s at 200-260°C and cutheats for 5 s at 20° below preheat temperature. Preheat plateaus are overall adequate. Recycling ratios are mostly within ±5% of unity, and the contribution of infrared stimulated luminescence to the OSL signal is <5%. Average D<sub>e</sub> and errors were calculated using the Central Age Model (Galbreith et al., 1999). “Aliquots used” are the number of aliquots used from those measured for calculating the average D<sub>e</sub>. Samples HR-7 and HR-16 are duplicates, taken from the same horizon 30 cm apart. Sample HR-11 is at the contact between 2 units and exhibit a bi-modal D<sub>e</sub> distribution; both modes are presented. Sample HR-14 was measured using the OSL and the ITL signals (180±11 ka and 183±13 ka, respectively). The D<sub>e</sub> and age presented here is the data combined from the two measurement protocols.